

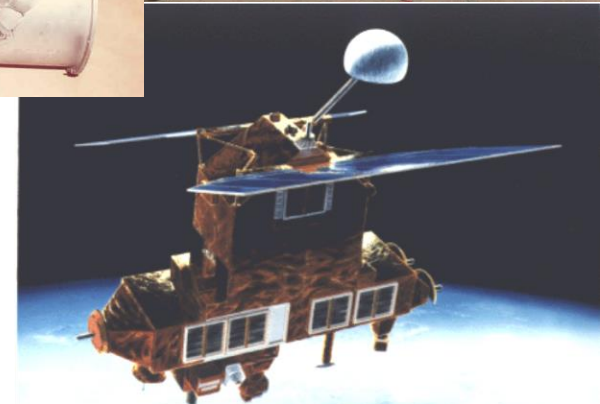
# **Modeling of the Orbital Debris Environment Risks in the Past, Present, and Future**

**Mark Matney, Ph.D.**  
**Orbital Debris Program Office**  
**NASA Johnson Space Center**



# Orbital Debris Types

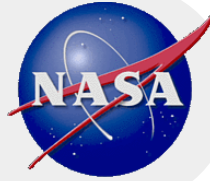
- **Intact objects, > 1 m**
  - Old rocket bodies
  - Spacecraft
  - “Operational” debris – shrouds, mounts, lens caps, etc
- **Fragmentation debris, 1 mm – 1 m**
  - **Deliberate or accidental explosions from on-board energy sources**
    - Unvented rocket fuel
    - Active batteries
    - Self-destruct mechanisms
  - **Deliberate or accidental collisions**
    - Weapons tests
    - Random collisions
  - **Solid rocket motor slag**
- **Degradation debris, < 1 mm**
  - **Deterioration of satellite surfaces in space environment**
    - Micrometeoroid and small debris impact ejecta
    - Paint deterioration in harsh space environment





# Modeling

- **NASA and the U.S. Dept. of Defense dedicate a tremendous amount of resources to measuring and monitoring the debris environment, but measurements do not always provide all the information we need**
  - Radars provide radar cross section (RCS), not size, material, shape, or mass
  - Similarly, optical telescopes provide brightness of reflected sunlight
  - **NASA uses a number of telescopes and radars to statistically sample only a subset of the environment**
    - Statistical sampling is the only way to measure objects <10 cm too small to track
  - **No matter how good or complete are our measurements, the orbital debris environment is dynamic. We cannot know with certainty what the environment will look like in the future**
- **The solution to these limitations is modeling**
  - **Modeling is the use of mathematical and compute tools to use the incomplete data we do have and determine the information we truly need**



## LEGEND

- One of NASA's "workhorse" models is **LEGEND**
- **LEGEND**, a LEO-to-GEO environment debris model
  - Is a high fidelity, three-dimensional numerical simulation model for long-term orbital debris evolutionary studies
  - Replaces the previous one-dimensional, LEO only model, EVOLVE
  - Includes intacts (R/Bs and S/C), mission-related debris (payload fairings, caps, etc.), and explosion/collision fragments
  - Handles objects individually
  - Is capable of simulating objects down to 1 mm in size, but the focus has been on  $\geq 10$  cm objects
  - Covers altitudes up to 40,000 km
  - Can project the environment several hundred years into the future
  - Uses a deterministic approach to "recreate" the historical debris environment based on recorded launches and breakups
  - Uses a Monte Carlo approach and a pair-wise collision probability evaluation algorithm to simulate future collision activities
  - Analyzes future debris environment based on user-specified launch traffic, post-mission disposal, and active debris removal options



# Similar OD Evolutionary Models

- **ASI's SDM**
  - Space Debris Mitigation long-term analysis program
- **ESA's DELTA**
  - Debris Environment Long-Term Analysis model
- **ISRO's KSCPROP**
  - Kustaanheimo and Stiefel Canonical Propagation model
- **JAXA's LEODEEM**
  - LEO Debris Environment Evolutionary Model
- **UKSA's DAMAGE**
  - Debris Analysis and Monitoring Architecture for the Geosynchronous Environment



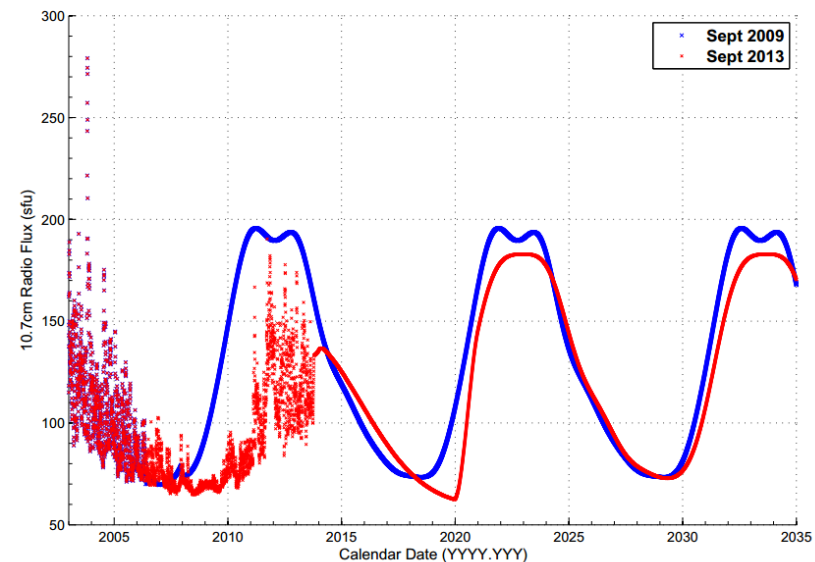
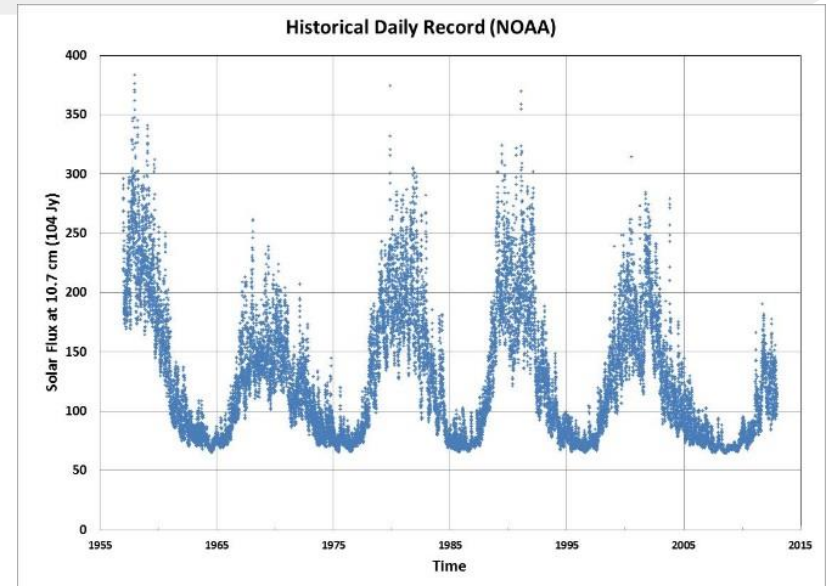
## LEGEND Supporting Models

- **LEGEND actually ties in many NASA models to do its calculations**
- **DBS database: a comprehensive record of historical launches and breakup events**
  - Time, type, orbit, physical properties (mass, area), *etc.*
  - The database is updated annually
- **U.S. Space Surveillance Network (SSN) catalogs**
  - Daily records of the historical growth of the  $\geq 10$  cm debris population
  - Orbit histories are used to derive empirical area-to-mass ratio ( $A/M$ ) distributions of breakup fragments
  - New files are downloaded from “Space Track” website daily
- **Future launch traffic model**
  - Typically a repeat of the last 8-year cycle, as commonly adopted by the international debris modeling community



# LEGEND Supporting Models

- **Atmospheric drag model**
  - Jacchia atmospheric density model (1977)
  - Drag perturbation equations based on King-Hele (1987)
  
- **Solar flux (at 10.7 cm wavelength) model consisting of three components**
  - Historical daily records available from the National Oceanic and Atmospheric Administration (NOAA) Space Weather Prediction Center (SWPC)
  - Short-term projection provided by NOAA/SWPC – currently through 2019
  - Long-term projection is a repeat of a 13th-order sine and cosine functional fit to Solar Cycles 18 to 24 (1944 – present)
    - Similar to projections developed for long-term debris evolutionary models by other space agencies (ASI, UKSA, etc.)





## LEGEND Supporting Models

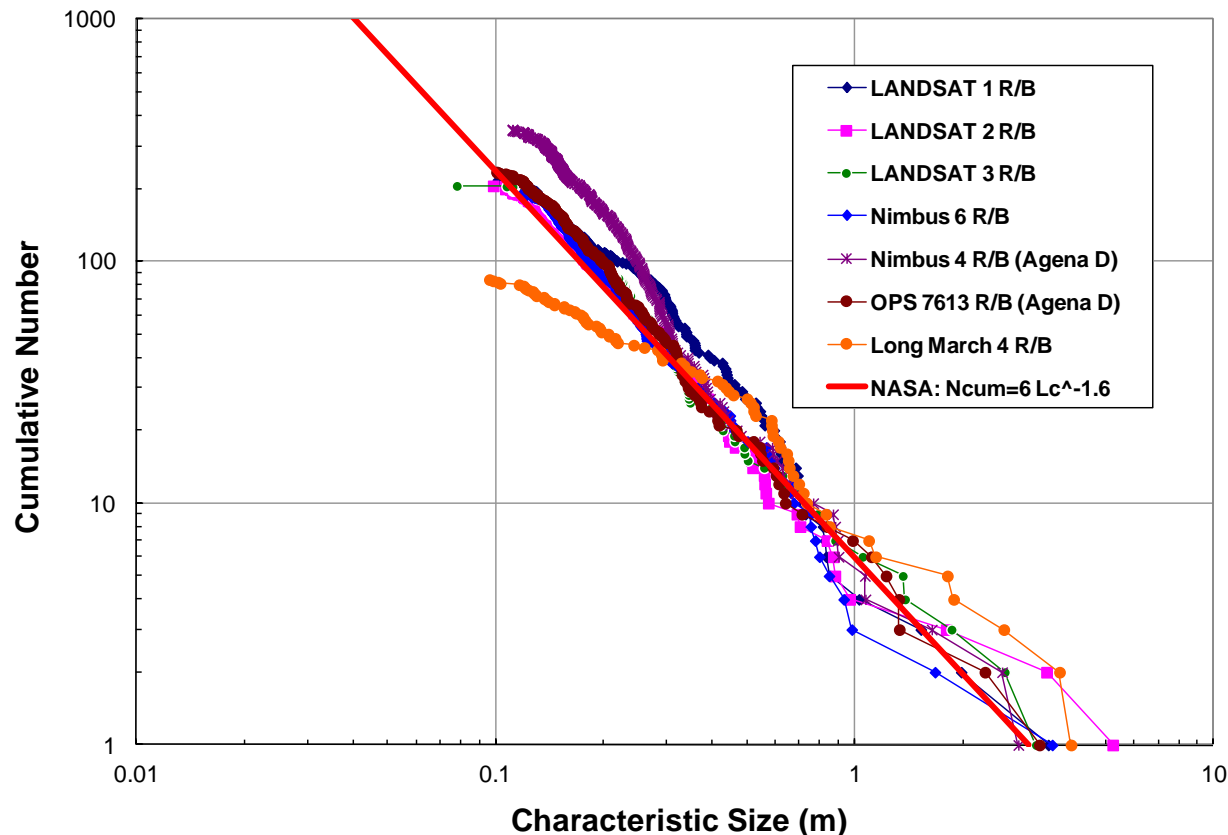
- **GEOprop orbital propagator**
  - Propagates objects near geosynchronous (GEO) region
  - Perturbations include solar and lunar gravitational forces, solar radiation pressure, and Earth's gravity-field zonal ( $J_2$ ,  $J_3$ , and  $J_4$ ) and tesseral ( $J_{2,2}$ ,  $J_{3,1}$ ,  $J_{3,3}$ ,  $J_{4,2}$ , and  $J_{4,4}$ ) harmonics
- **Prop3D orbit propagator**
  - Propagates orbits of objects in LEO and GTO regions
  - Perturbations include atmospheric drag, solar and lunar gravitational forces, solar radiation pressure, and Earth's gravity-field zonal harmonics  $J_2$ ,  $J_3$ , and  $J_4$
- **Both propagators compare reasonably well with the evolution of the SSN cataloged objects**





# LEGEND Supporting Models

- **NASA Standard Satellite Breakup Model**
  - Describes the outcome of an explosion or collision
    - Fragment size, A/M, and  $\Delta V$  distributions
  - Based on seven, well-observed on-orbit explosions, several ground-based impact experiments, and one on-orbit collision





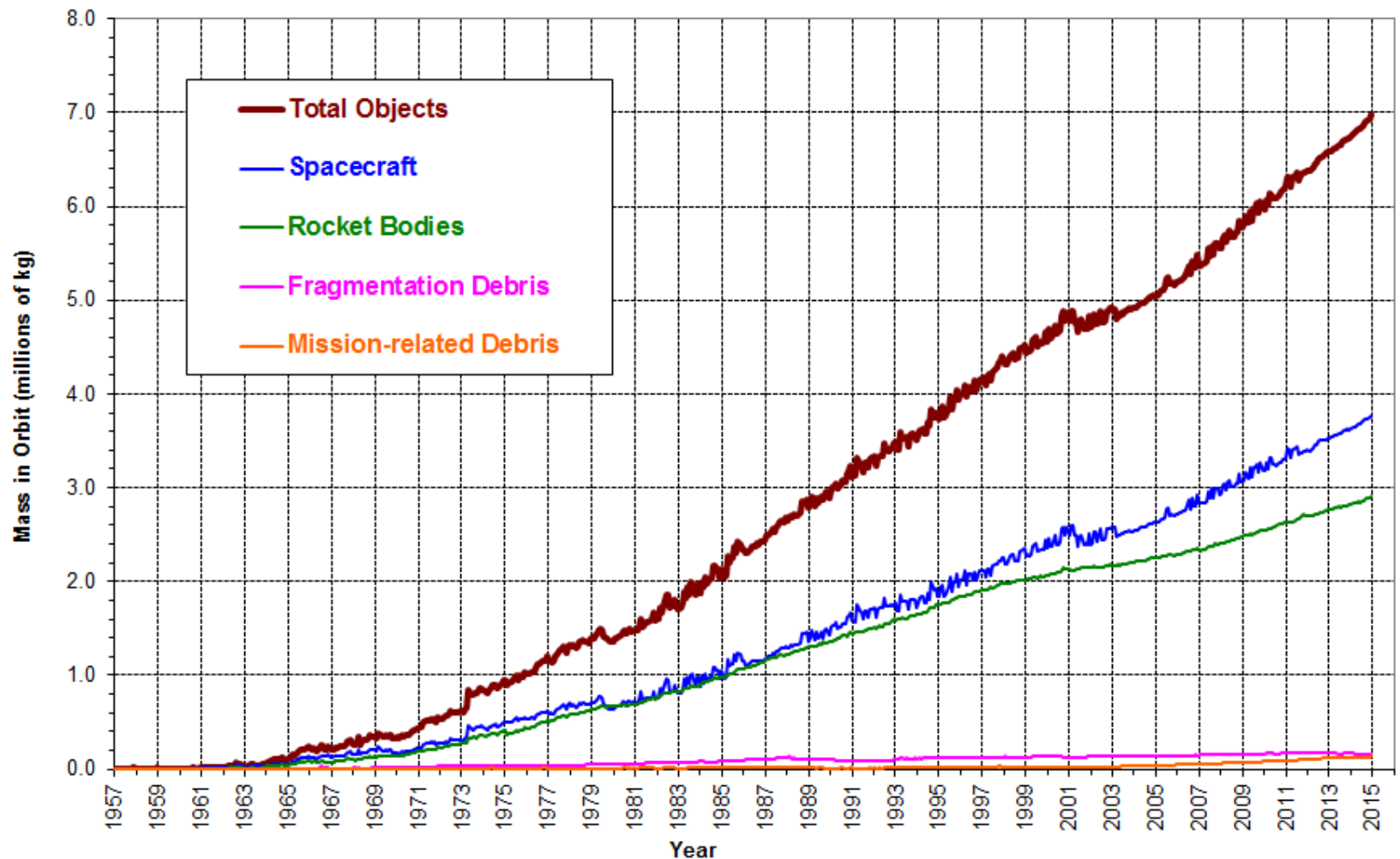
# LEGEND Applications

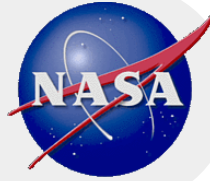
- **LEGEND is the tool the NASA Orbital Debris Program Office uses to**
  - **Provide debris environment projection for the next 200 years**
    - Based on user-specified scenarios (launch traffics, postmission disposal, active debris removal options, etc)
  - **Evaluate the instability of the current debris environment**
  - **Assess the growth of the future debris populations**
  - **Characterize the effectiveness of the NASA, U.S., and international debris mitigation measures**
  - **Quantify the benefits of active debris removal (ADR)**



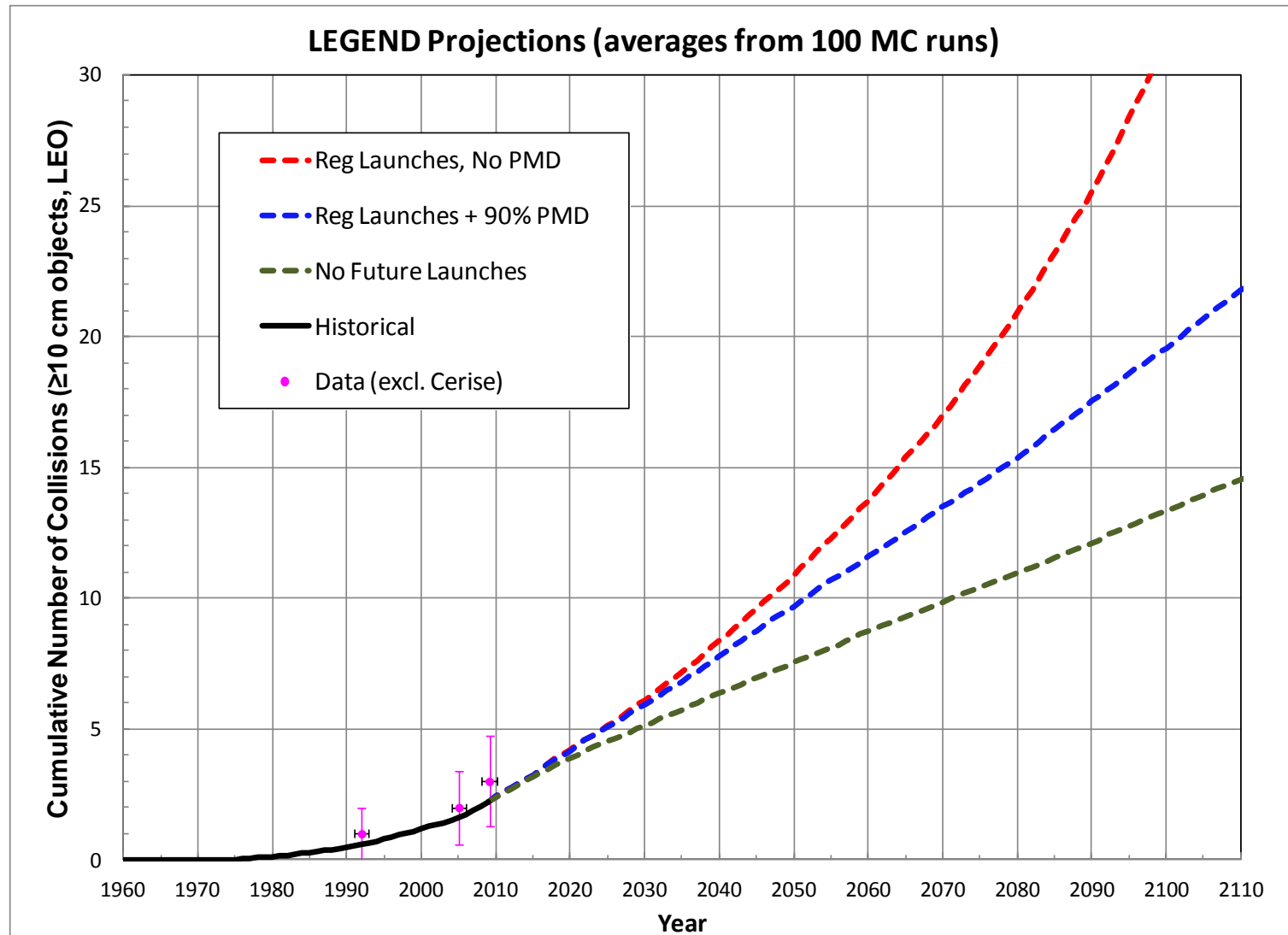
# Mass Accumulation in Orbit – Based on DBS

Monthly Mass of Objects in Earth Orbit by Object Type





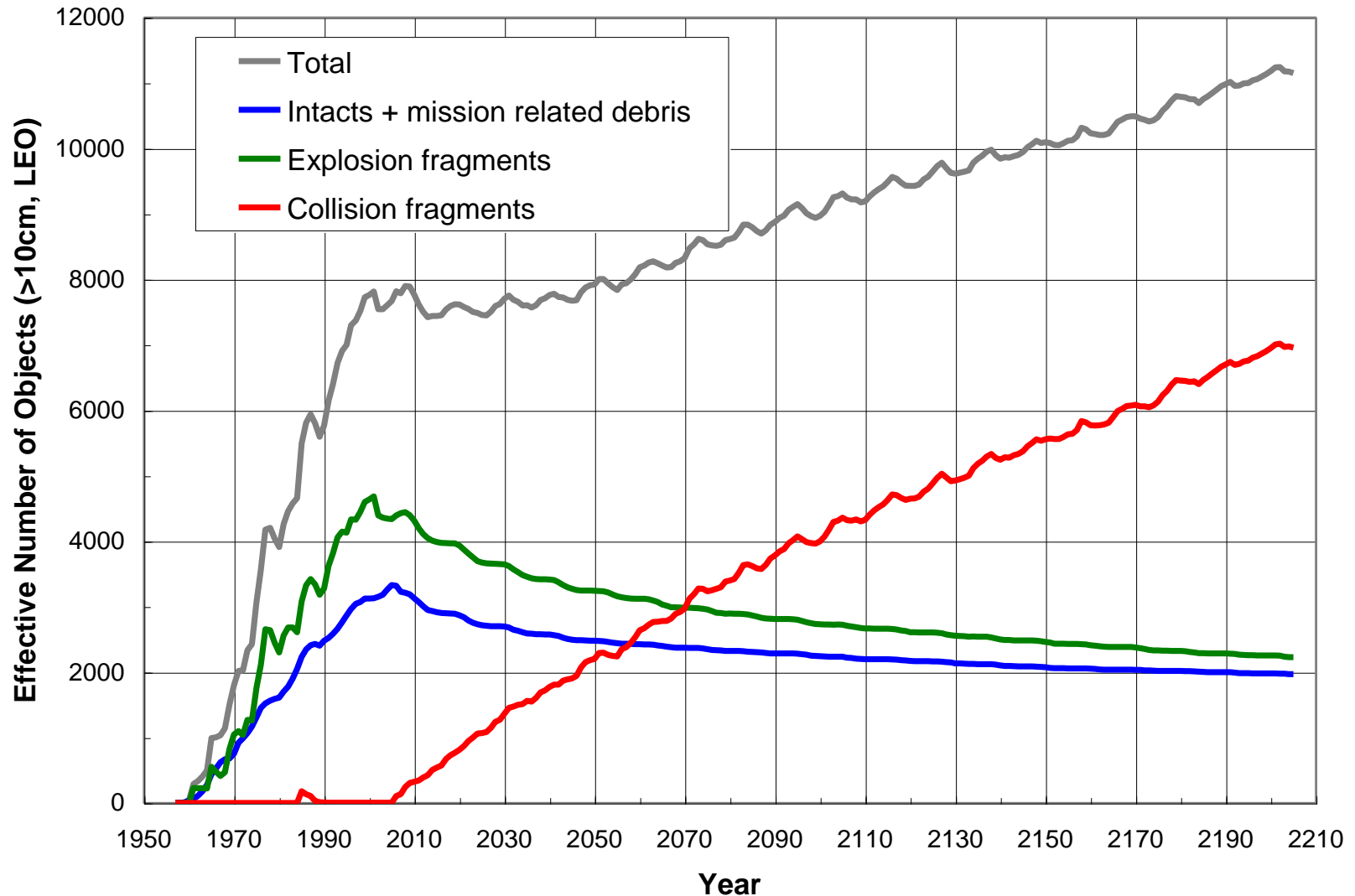
# Sample LEGEND Output – Collisions in LEO

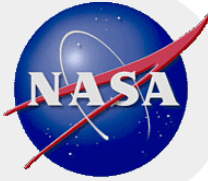




# Growth with no future launches

## Kessler Syndrome





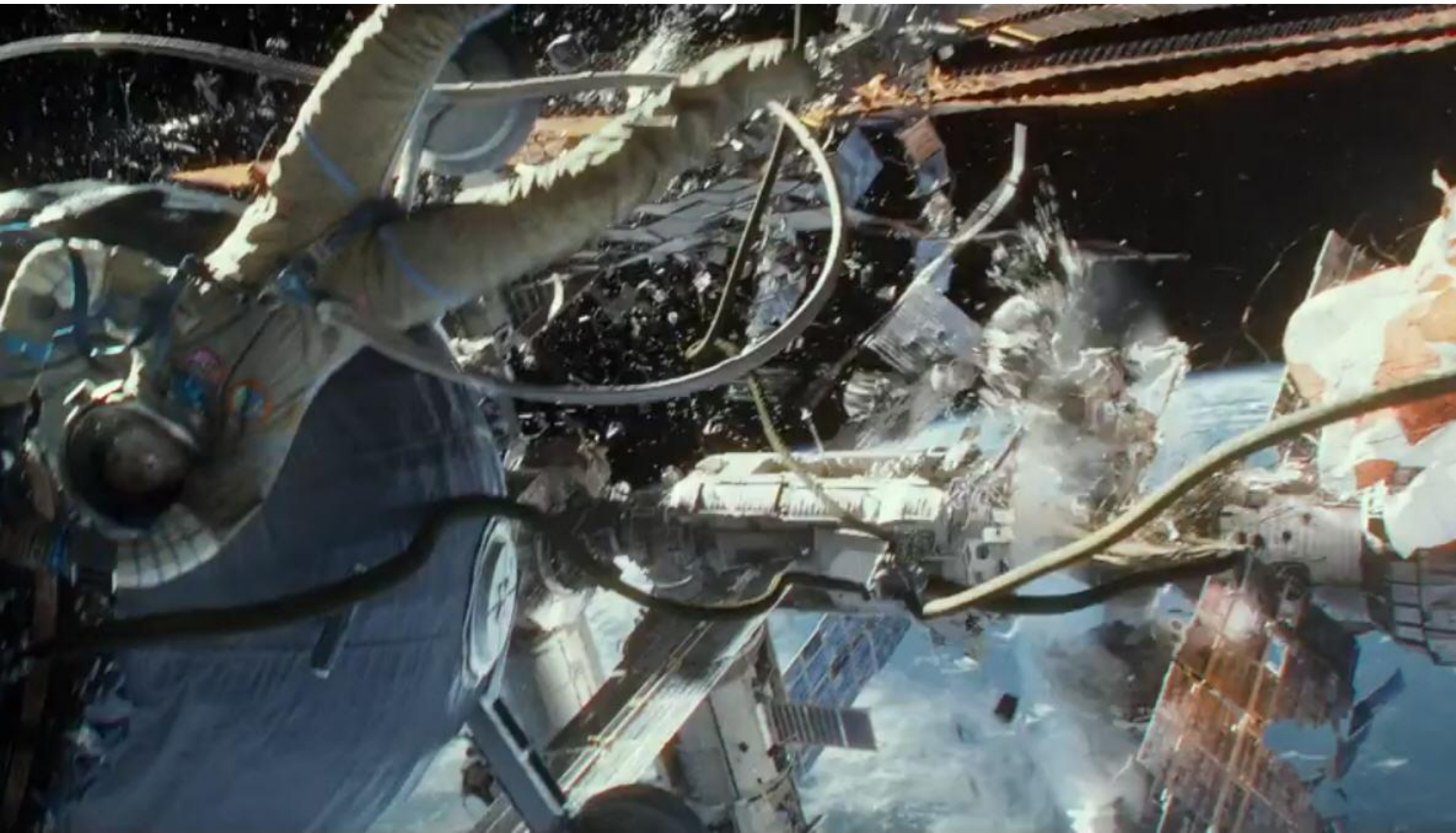
## Rubes



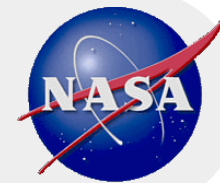
"Well, I'll be ... I guess the little chicken was right."



# Gravity







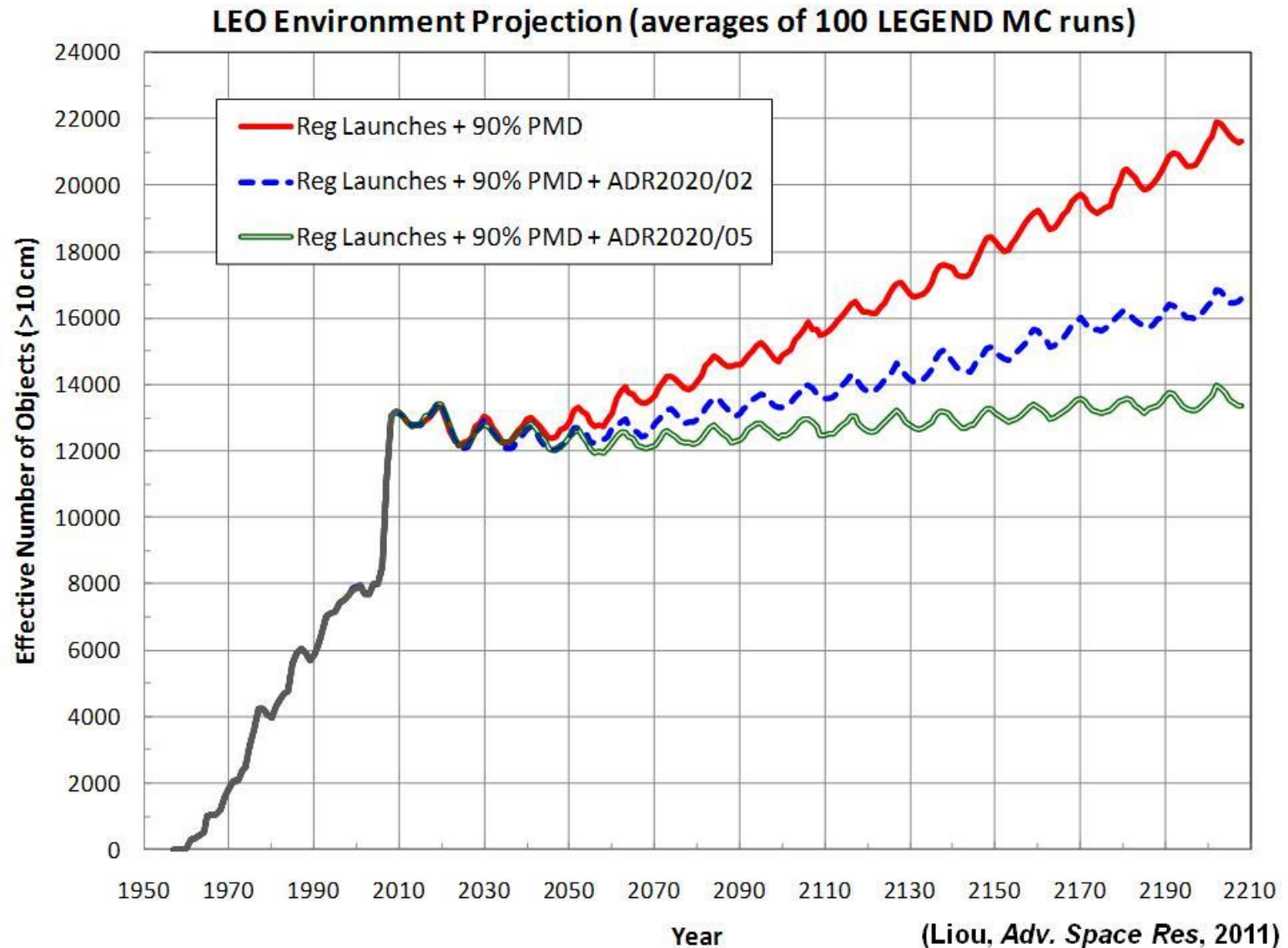
# Gravity

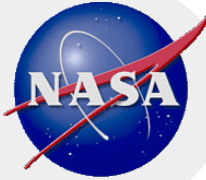






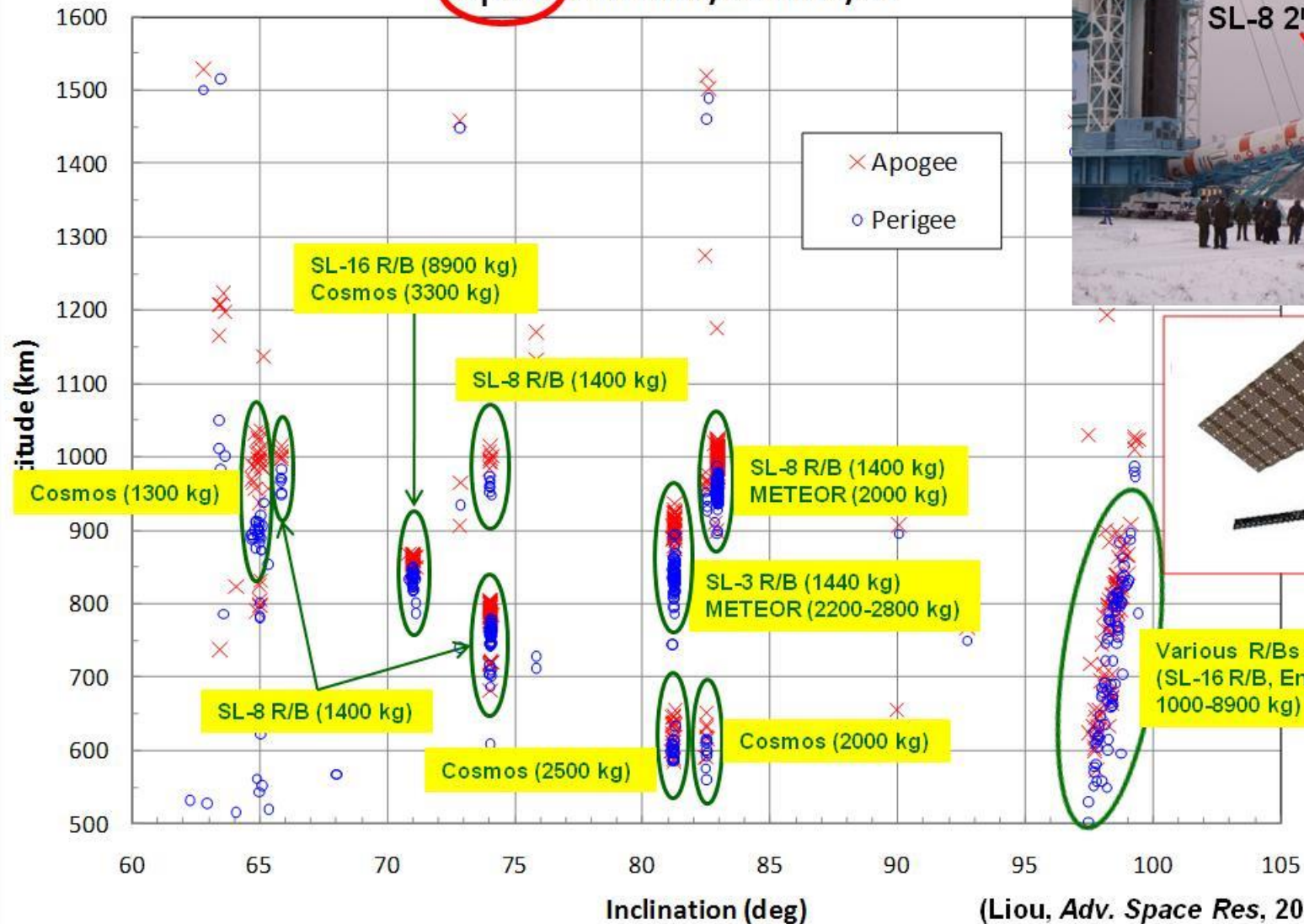
## Fix the Problem? – Remove Mass





# Highest Mass Objects

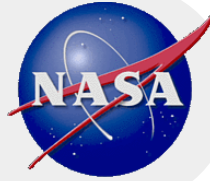
**Top 500 Current R/Bs and S/Cs**





## Active Debris Removal Cartoon, 1965 (!)





## ORDEM 3.0

- **An Engineering Model is a tool (primarily) for spacecraft designers and users to understand the long-term risks of debris collisions with their spacecraft**
- **NASA's Orbital Debris Engineering Model ORDEM 3.0 represents NASA's best estimate of the current and near future orbital debris environment.**
  - The environment is dynamic and must be updated periodically
- **ORDEM 3.0 has significant new capabilities over past ORDEM models**
  - Uncertainties
  - Material density categories
  - Model extended to GEO
  - Can easily calculate flux for satellites in highly elliptical orbit



## ORDEM 3.0 vs. ORDEM2000

Parameter	ORDEM2000	ORDEM 3.0
Spacecraft & telescope/radar analysis modes	Yes	Yes
Time range	1991 to 2030	2010 to 2035
Altitude range with minimum debris size	200 to 2000 km (>10 $\mu\text{m}$ ) (LEO )	200 to 38,000 km (>10 $\mu\text{m}$ ) (LEO to GTO) 34,000 to 38,000 km (>10 cm) (GEO)
Orbit types	Circular (radial velocity ignored)	Circular to highly elliptical
Model populations divided by type & material density	No	Intacts Low-density (<2 g/cc) – e.g., plastic Medium-density (2-6 g/cc) – e.g., aluminum High-density (>6 g/cc) – e.g., steel RORSAT NaK coolant droplets (0.9 g/cc)
Special model populations	No	Yes (ASAT, Iridium/Cosmos, Snapshot, Transit)
Model cumulative size thresholds ( <i>fiducial points</i> )	10 $\mu\text{m}$ , 100 $\mu\text{m}$ , 1 mm, 1 cm , 10 cm, 1 m	10 $\mu\text{m}$ , 31.6 $\mu\text{m}$ , 100 $\mu\text{m}$ , 316 $\mu\text{m}$ , 1 mm, 3.16 mm, 1 cm, 3.16 cm, 10 cm, 31.6 cm, 1 m
Flux uncertainties	No	Yes
Meteoroids	No	No*

\* a separate meteoroid environment model (MEM) is available from NASA's Meteoroid Environment Office



# **ORDEM Process**

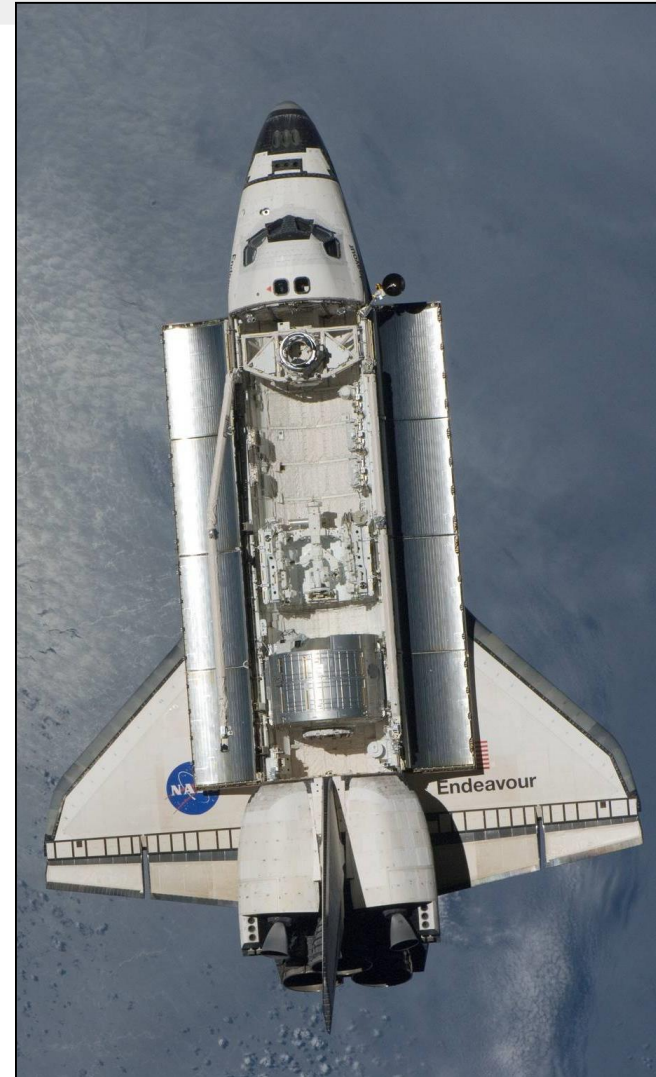
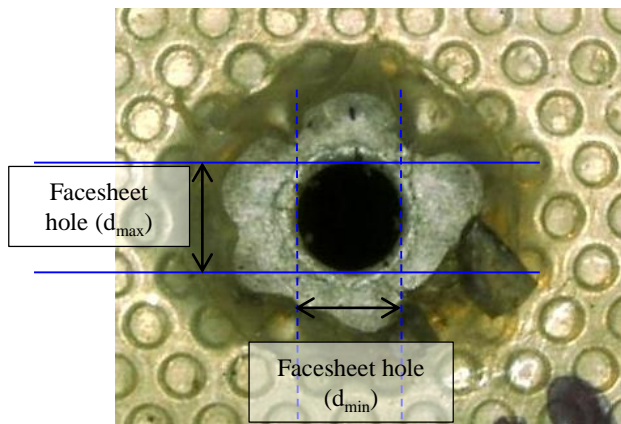
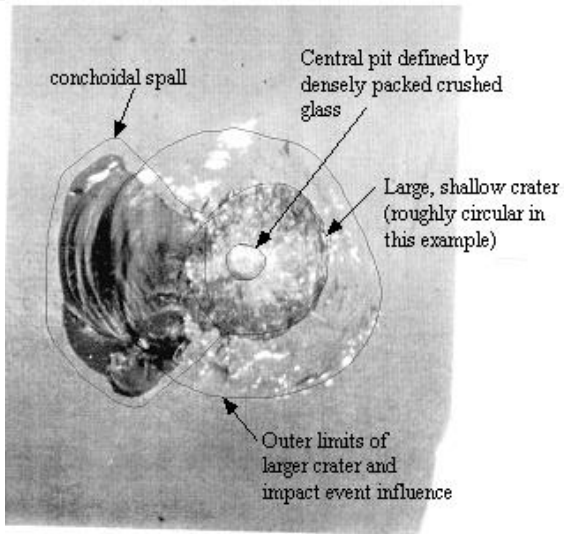
## **Creating the Current Environment**

- **Initial environment created by using database of known space activity and tools such as the NASA Standard Breakup Model (to model breakup clouds) & PROP3D (to model long-term orbit evolution)**
- **Environment-dominating events such as the Chinese ASAT (~850 km) and the Iridium/COSMOS collision (~775 km) were modeled separately as were a few unique non-breakup populations**
- **Debris material densities**
  - **For sub-mm debris - determined from analysis of residue in impact features from returned spacecraft surfaces (specifically, Shuttle windows and radiators)**
  - **For larger debris - directly measured from ground impact tests**
- **Maximum Likelihood Estimator used to empirically fit the environment to measurement data, creating a final “Current” debris environment**
  - **This resulted in adjusting model populations to fit data using size-dependent “weighting factors”**
  - **Size-dependent weighting factors derived from these data-fitting processes are used to project into the future**
  - **Uncertainties computed using Bayesian and other techniques**





# Shuttle *In Situ* Data





# ORDEM 3.0 Datasets and Supporting Models

Observational Data	Role	Region/Size
SSN catalog (radars, telescopes)	Intacts & large fragments	LEO > 10 cm, GEO > 70 cm
HAX (radar)	Statistical populations	LEO > 3 cm
Haystack (radar)	Statistical populations	LEO > 5.5 mm
Goldstone (radar)	Statistical populations	LEO > 3 mm
STS windows & radiators (returned surfaces)	Statistical populations	$10 \mu\text{m} < \text{LEO} \leq 1 \text{ mm}$
MODEST (telescope)	GEO data set	GEO > 30 cm

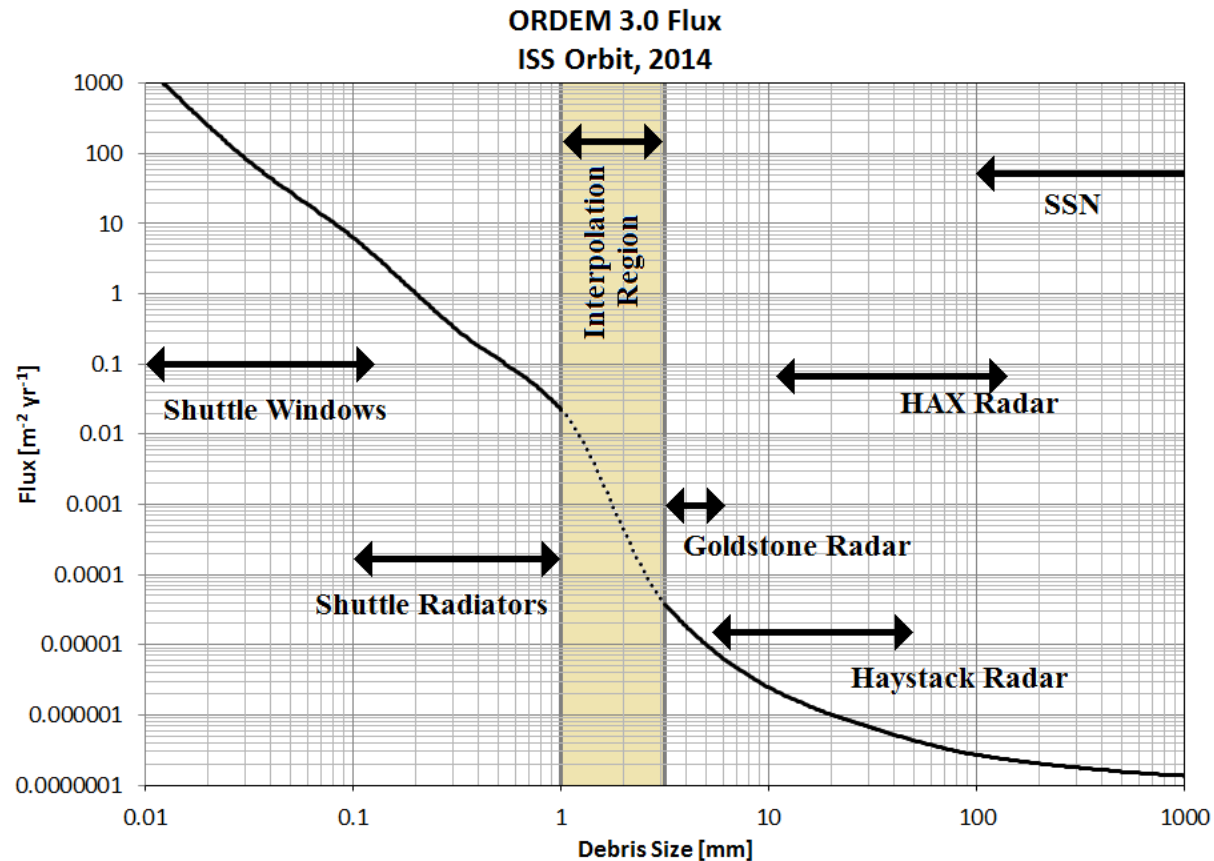
- Note that the US Space Shuttle is no longer an active data source







# Data and Size Regimes



- Small particle populations are fit separately from large particle populations



## ORDEM

# Projecting Into the Future – Debris > 3 mm

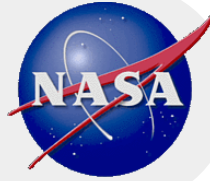
- **LEGEND** used to propagate the “Current” environment into the future
- Populations empirically adjusted to match radar and optical measurement data
- When **LEGEND** creates new future debris (through future collisions or explosions) the same weighting values that were used to fit historical size distributions are applied to debris production in the future
- Launch rate, solar activity, and explosion rate are independent inputs into the model
- 120 Monte Carlo future environments are created
  - Future collisions simulated stochastically
- Reported future environment is the average of the 120 possible future environments, “spread” of possible futures preserved as population uncertainties



## ORDEM

### Projecting Into the Future – Debris < 1 mm

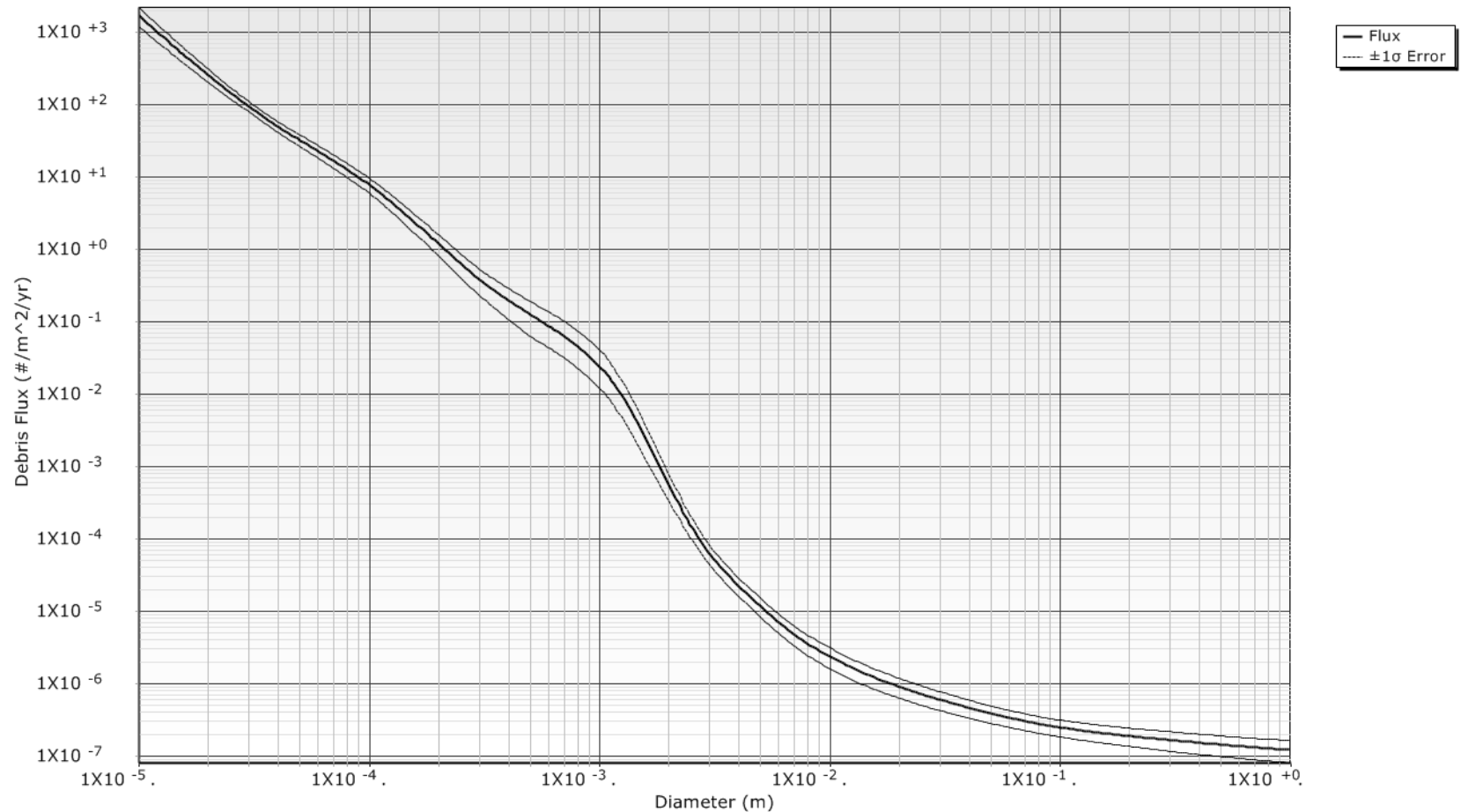
- **LEGEND** used to characterize the population of intact objects in the future as source objects for small debris
- The surface degradation model “creates” particles with zero delta-velocity at different sizes and material types proportional to the area of the parent body
- These debris are propagated under solar radiation pressure and atmospheric drag to compute flux on *in situ* surfaces
- Damage equations (based on empirical tests) are used to “predict” distribution in feature size (e.g., crater diameter) on the *in situ* surface using reference debris population
- Production rates at the parent bodies adjusted to match empirical data



# ORDEM Flux for ISS 400km

## Average Cross-Sectional Flux vs. Size

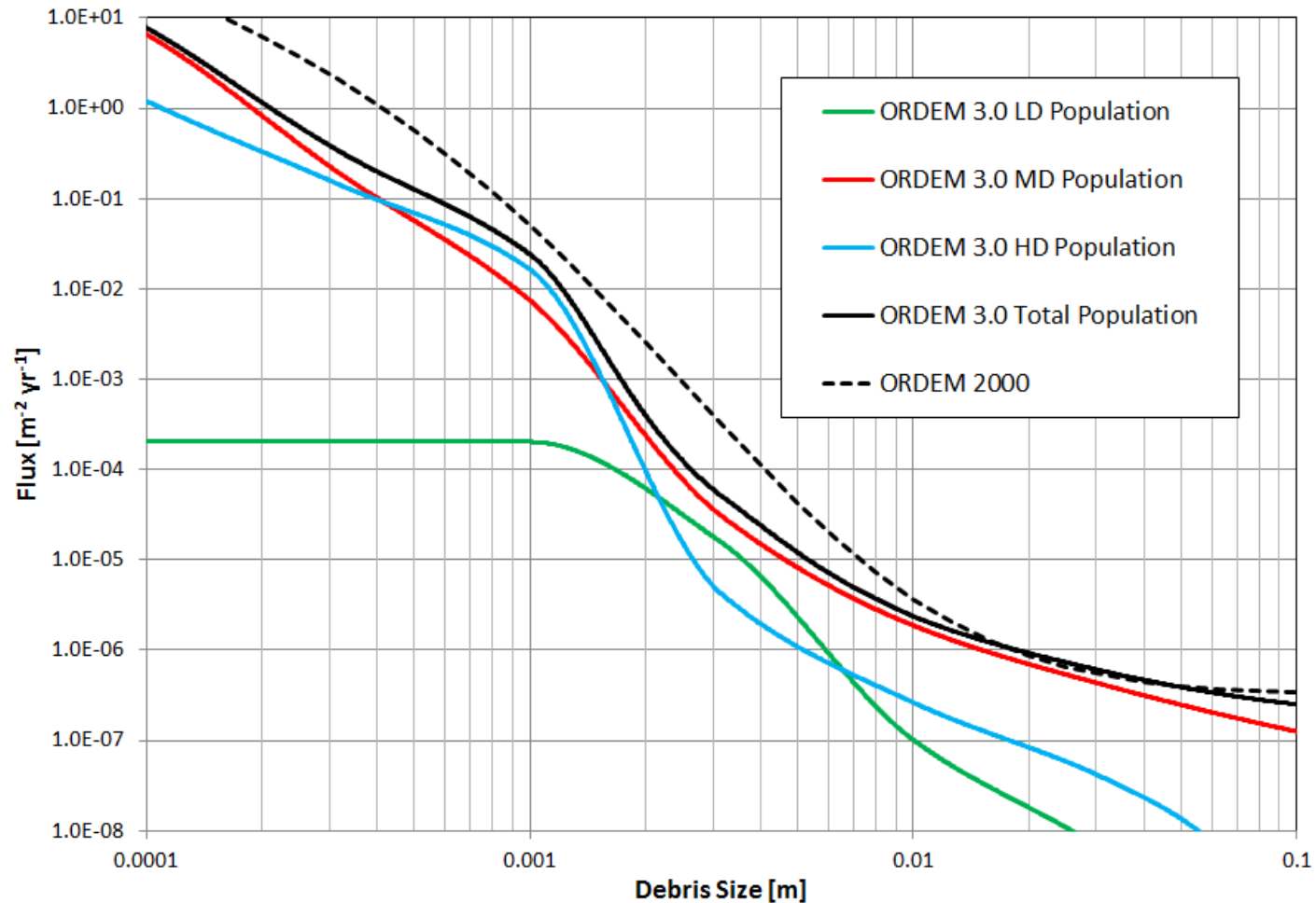
Year: 2013 Perigee Altitude = 400.000 Apogee Altitude = 400.000 inc = 51.60





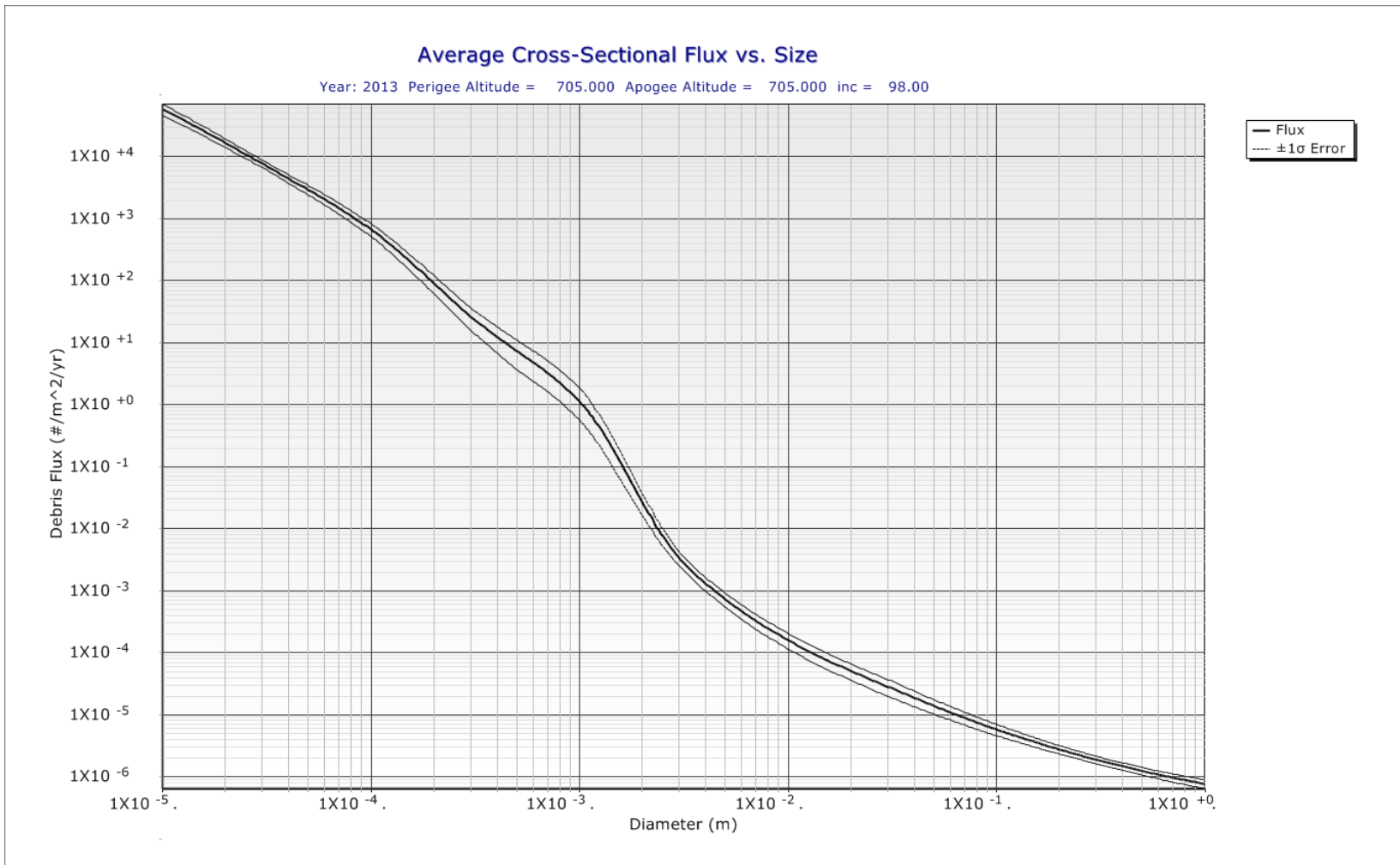
# Material Distributions - ISS

ORDEM Populations for 2013 ISS Flux as a Function of Debris Size





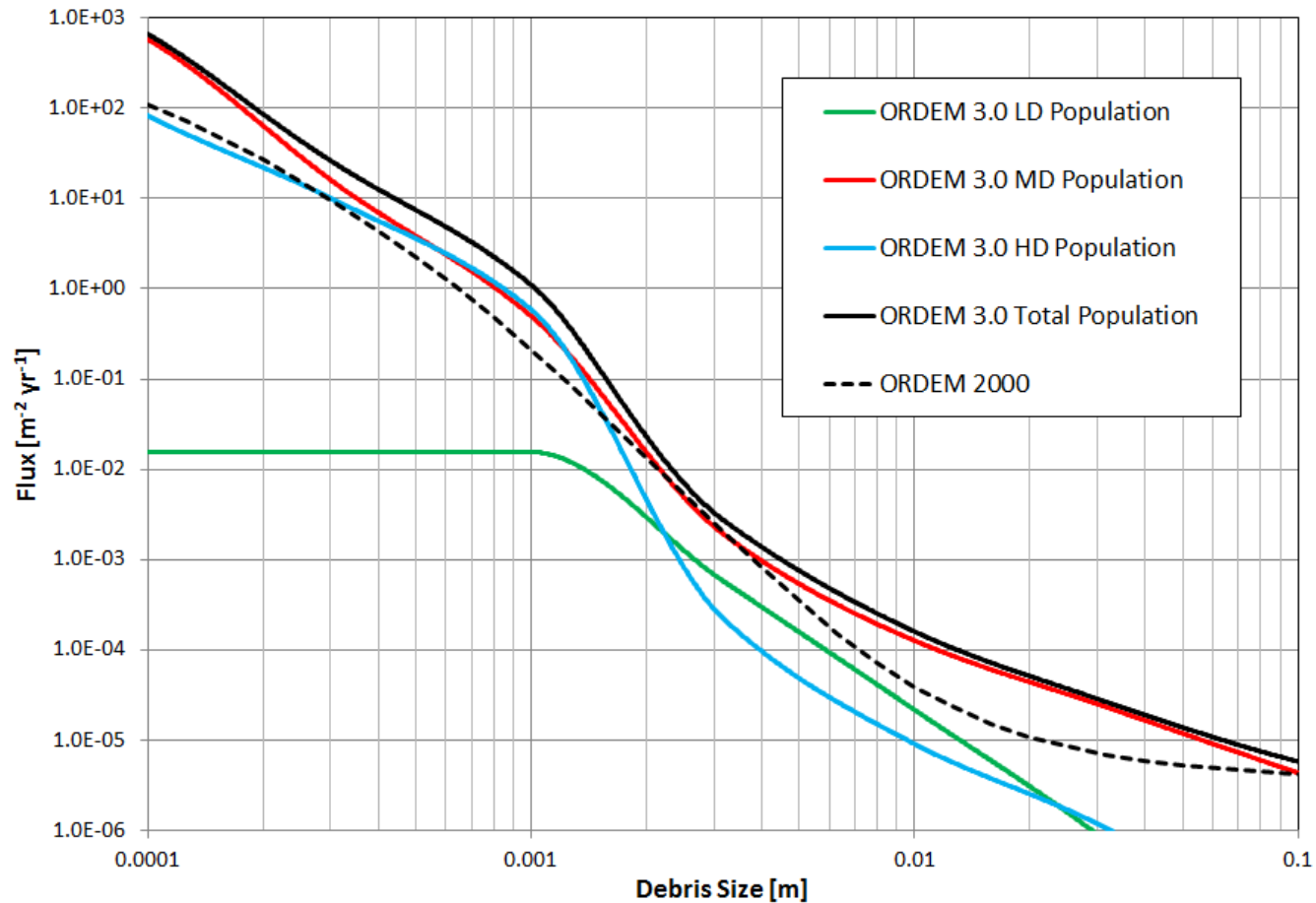
# ORDEM Flux for A-Train 705km



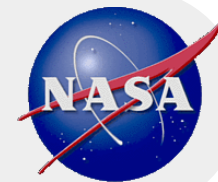


# Material Distribution – A-Train

ORDEM Populations for 2013 98° 705 km Orbit Flux as a Function of Debris Size



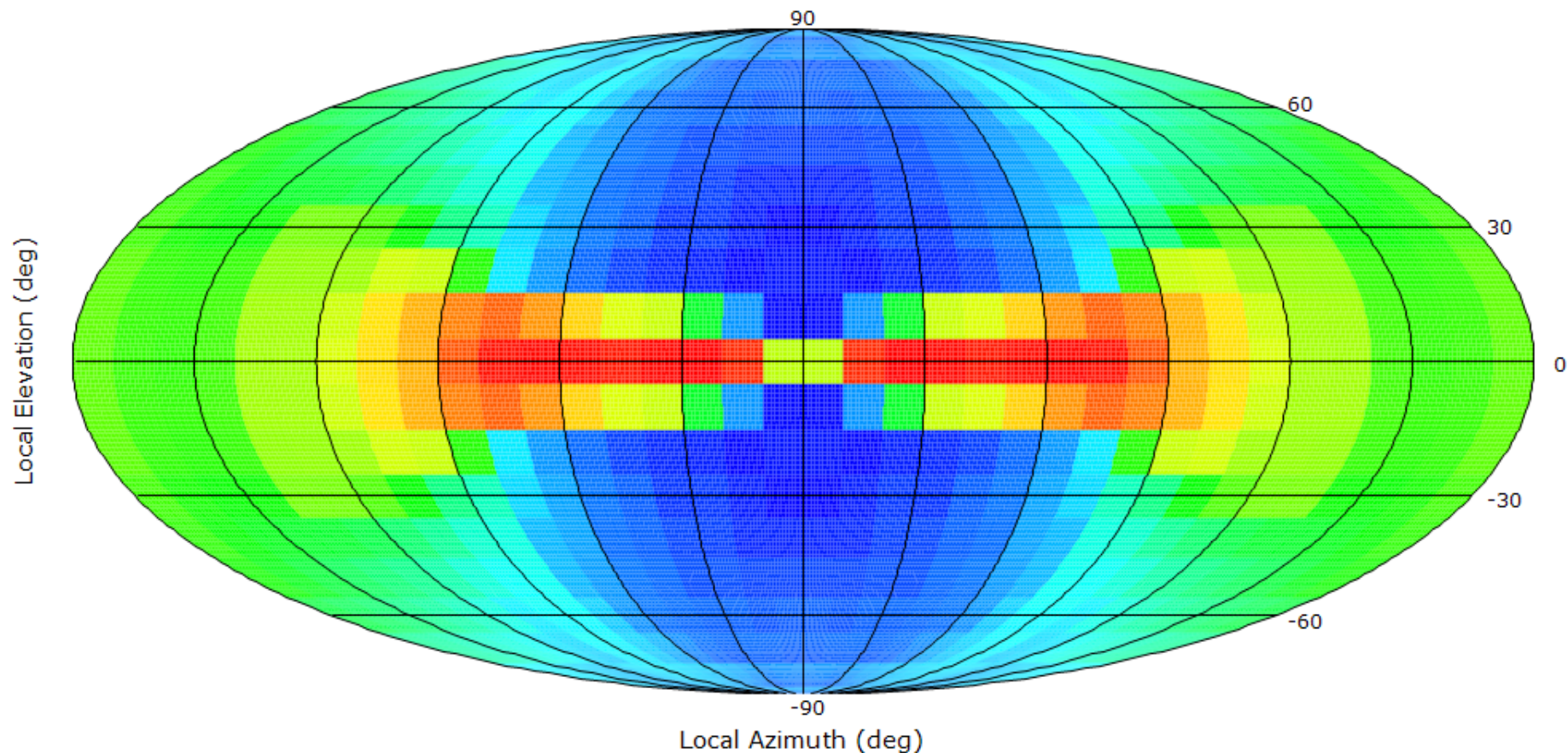




# ORDEM 3.0 Outputs

## 2-D Directional Flux

Year: 2013 Perigee Altitude = 400.000 Apogee Altitude = 400.000 inc = 51.60 particle size = >1mm



7/23/2013 5:06:05 PM

$10^{-12}$

$10^{-4}$

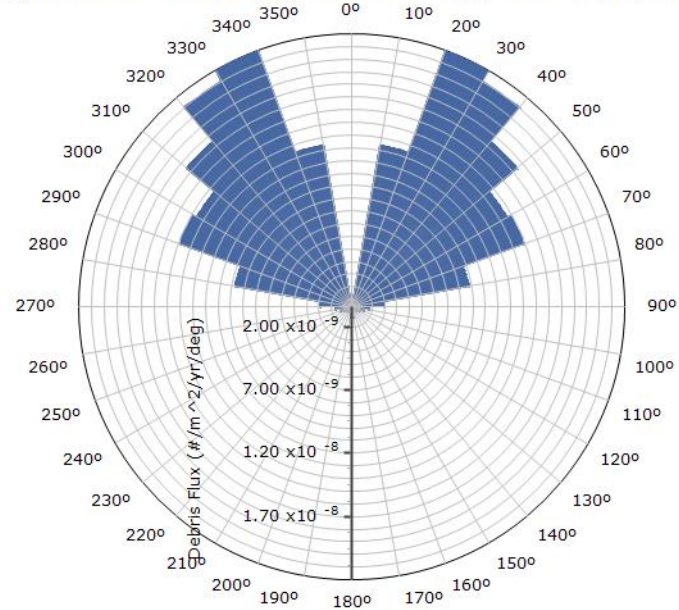




# ORDEM 3.0 Outputs

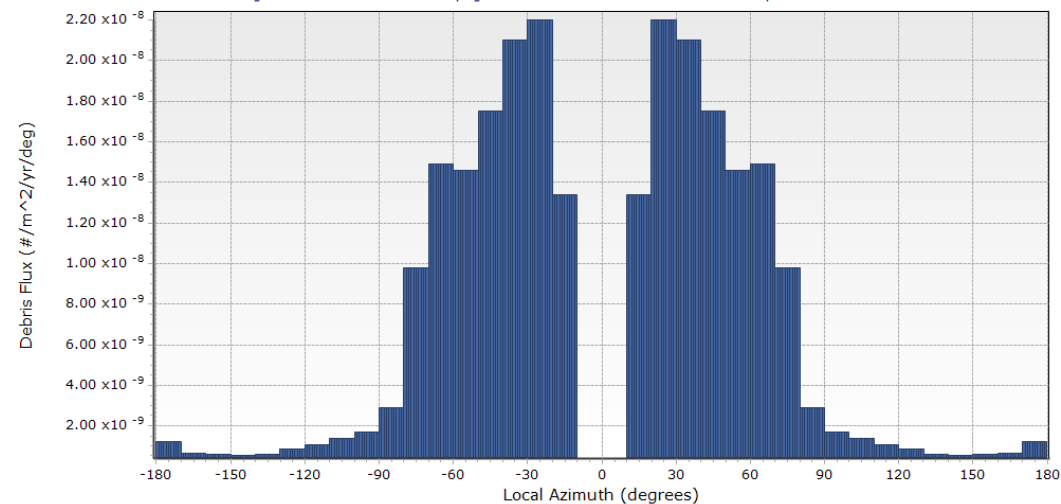
## Flux vs. Local Azimuth

Year: 2013 Perigee Altitude = 400.000 Apogee Altitude = 400.000 inc = 51.60 particle size = >1cm



## Flux vs. Local Azimuth

Year: 2013 Perigee Altitude = 400.000 Apogee Altitude = 400.000 inc = 51.60 particle size = >1cm



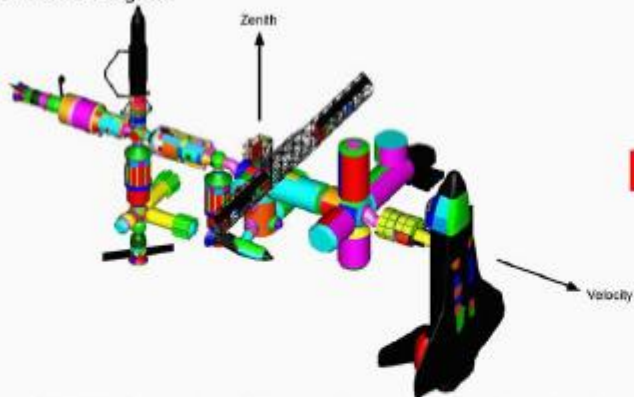


# BUMPER

## NASA/JSC BUMPER-II Meteoroid/Debris Threat Assessment Code

### Spacecraft Configuration (I-DEAS Finite Element Model)

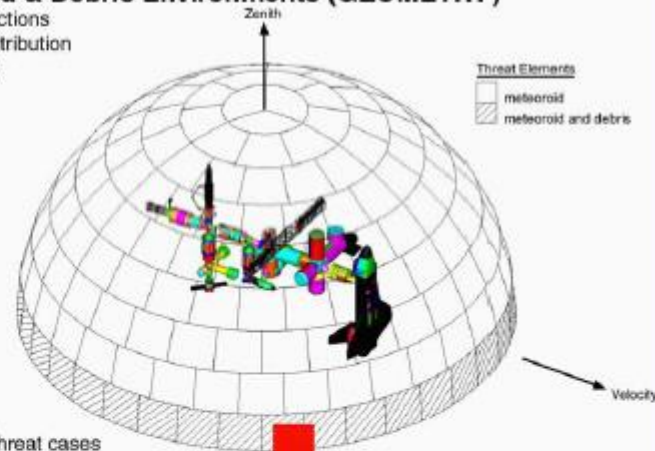
- Describes spatial relationships of spacecraft components
- Defines spacecraft orientation (velocity and zenith directions)
- Defines M/OD shield regions



- Approximately 120,000 elements in ISS assembly complete mated configuration FEM

### Meteoroid & Debris Environments (GEOMETRY)

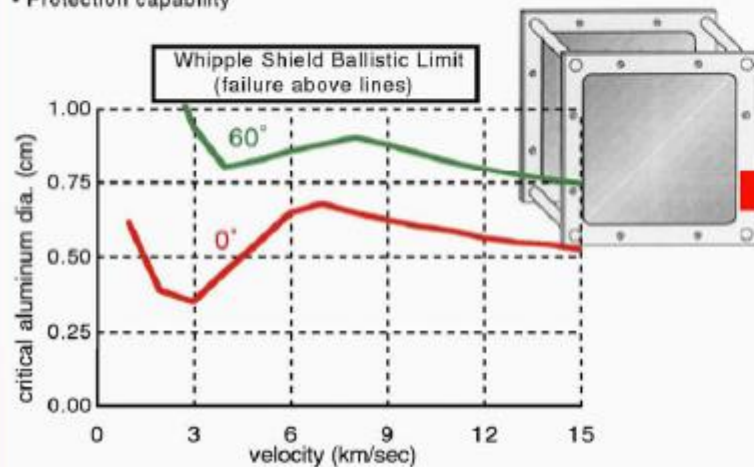
- Threat directions
- Velocity distribution
- Shadowing



- 90 debris threat cases and 149 meteoroid threat cases assessed for each element in the FEM

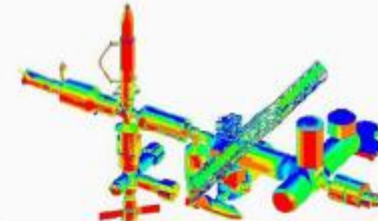
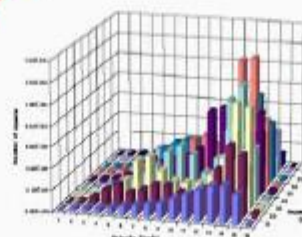
### Critical Particle Diameter Calculation (RESPONSE)

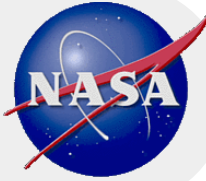
- Protection capability



### Computation of Penetrating Flux and PNP (SHIELD) Graphical Interpretation of Results (EXCEL & I-DEAS)

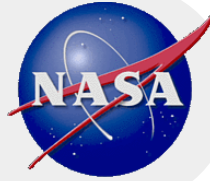
Station Region	Impact Risk From 1 mm Q Debris		Debris Penetration Risk	
	Probability No Impact	Odds of Impact	Probability No Penetration	Odds of Penetration
FCB	0.996536	1/214	0.996541	1/224
Service Module	0.999336	1/1506	0.999796	1/4912
Node 2	0.990466	1/106	0.999996	1/625000
Hab Module	0.996074	1/29	0.999923	1/928
Lab Module	0.996532	1/69	0.999922	1/1029
GRV	0.997446	1/61	0.999930	1/6220
TOTALS	0.996822	1/15	0.996132	1/146





## Collision Avoidance

- **Orbiting objects larger than about 10 cm are tracked by the U.S. Dept. of Defense (DoD) Space Surveillance Network (SSN)**
- **Current statistical technique was developed as a joint project by the DoD and NASA to ensure the safety of Shuttle and ISS astronauts**
- **Service now provided to any space user**
  - **Possible conjunction warning given to registered user**
  - **Contains the covariance matrix and encounter geometry for each object**
    - Covariance matrix gives uncertainty ellipsoid of the position of each object
    - Information can be used to compute a probability of collision
- **For every object tracked, there are tens to hundreds of objects we cannot track that can still cause serious damage to a spacecraft**
  - **Collision avoidance is prudent, but does not solve the debris problem**
- **Vast majority of objects tracked (~95%) are inert and cannot maneuver**
  - **Not a solution for problem of long-term collisional growth**



# Reentry Modeling

- **NASA's ORSAT code is used to assess the survivability of reentering objects in order to account for risk to ground population**
- **Cases are run in hierarchy beginning with parent body and then moving on to components and sub-components if necessary**
- **User makes decisions on how components are modeled**
  - Breakup altitude
  - Component shape selection
  - Component motion
  - When component heating begins
- **Options exist to run parametric study on specific variables:**
  - oxidation efficiency
  - initial wall temperature
  - surface emissivity
  - breakup altitude



# ORSAT Code Structure

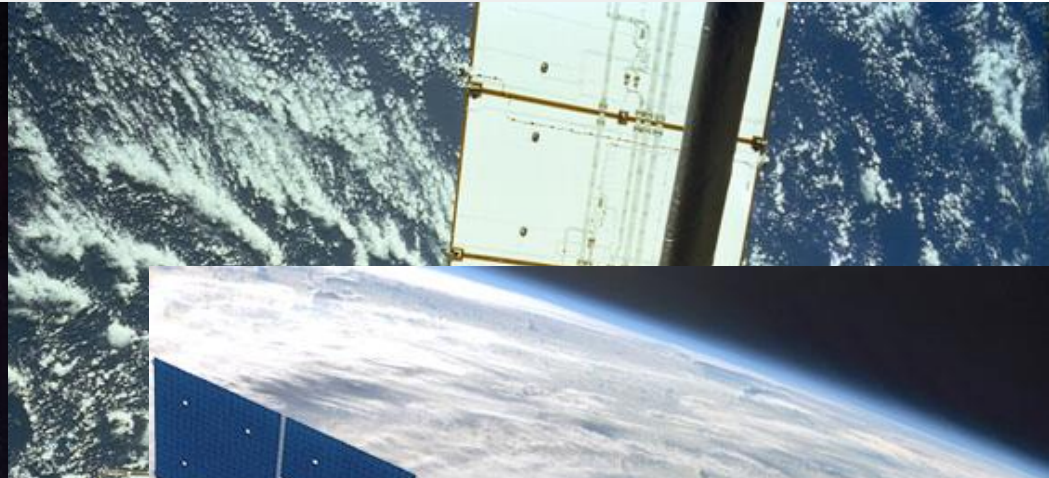
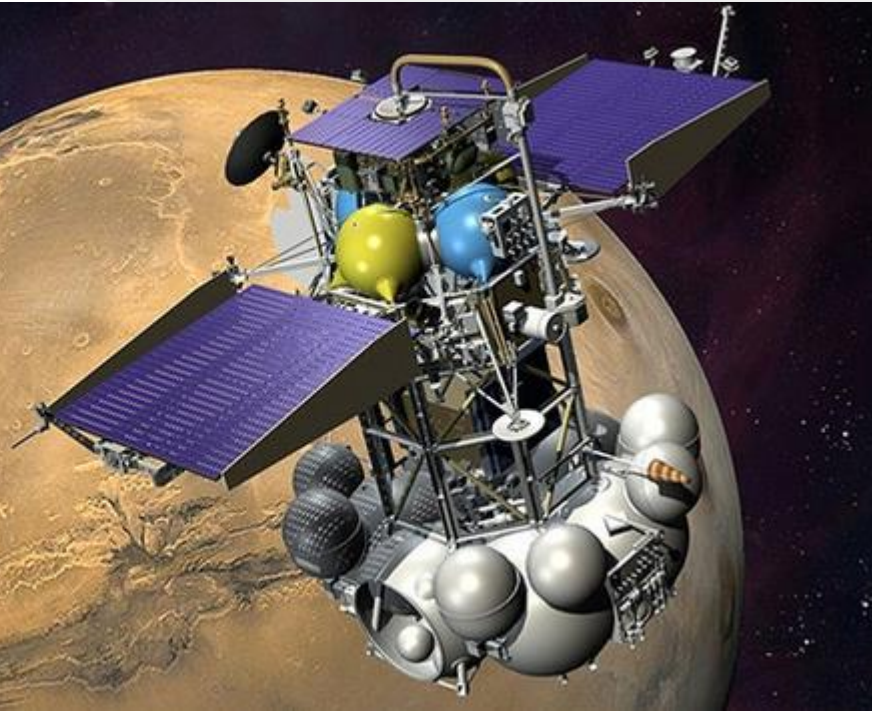
- **Six general modules in code:**
  - **Trajectory**
    - Recently updated in ORSAT 6.0 and 6.1 from Miehle method to Vinh
  - **Atmosphere**
    - Chosen from 3 models or input a user-defined model
    - Difference between 3 models is small due to only small changes in density
  - **Aerodynamics**
    - Recent updates to low altitude/Mach number drag coefficients
  - **Aerothermodynamics**
    - Detra-Kemp-Riddell or Fay-Riddell (small variations in results)
    - Stagnation point heating is well developed
    - Averages and 3D distributions over various shapes
  - **Thermal**
    - Different modes can lead to different results
  - **Debris Casualty Area / Risk**





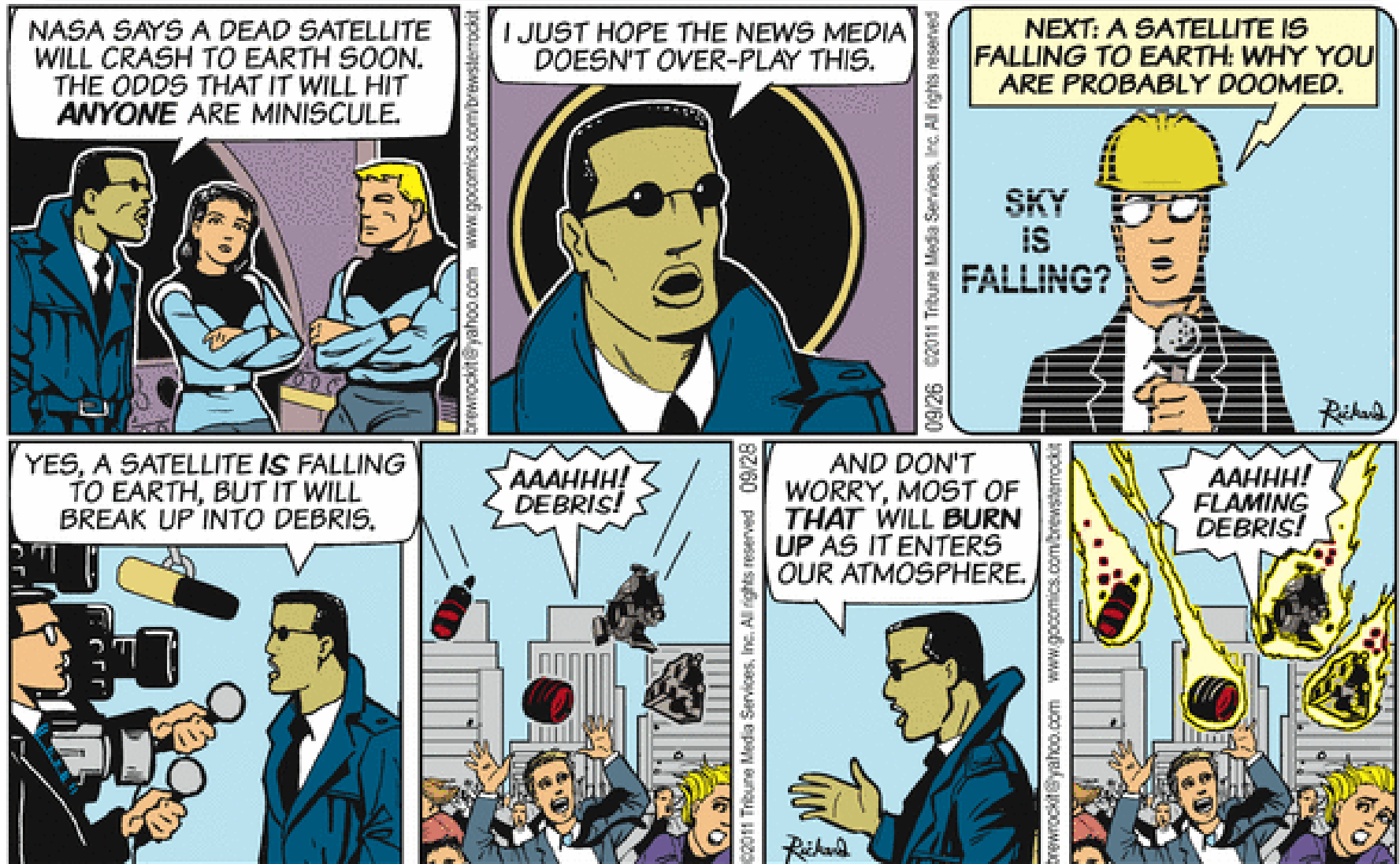
# Recent Reentries

## UARS, ROSAT, Phobos-Grunt, TRMM

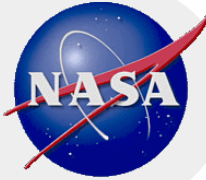




# UARS Reentry in the Popular Imagination

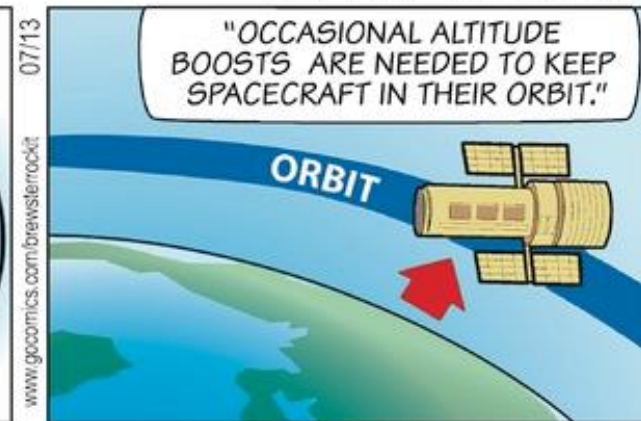
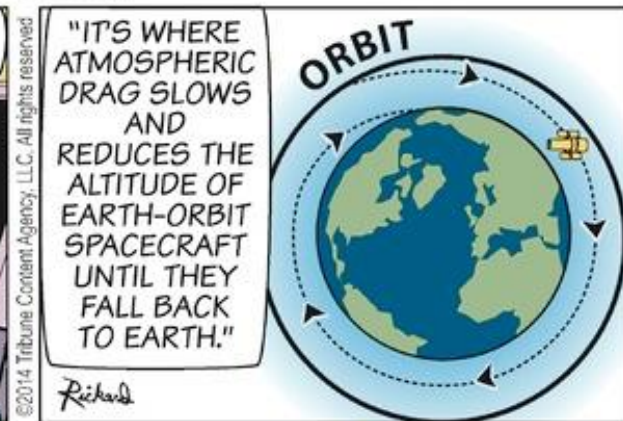
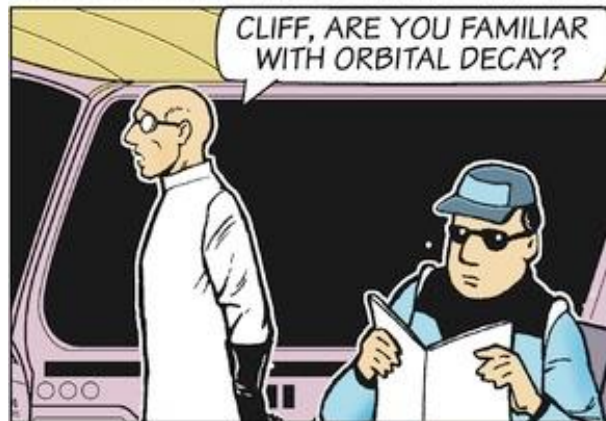






## BREWSTER ROCKIT: SPACE GUY!

BY TIM RICKARD







## That Which Survives...



**Texas, 1997**



**South Africa, 2000**



**Zimbabwe, 2013**



**Guatemala, 2003**



**Argentina, 2004**



**Saudi Arabia, 2001**



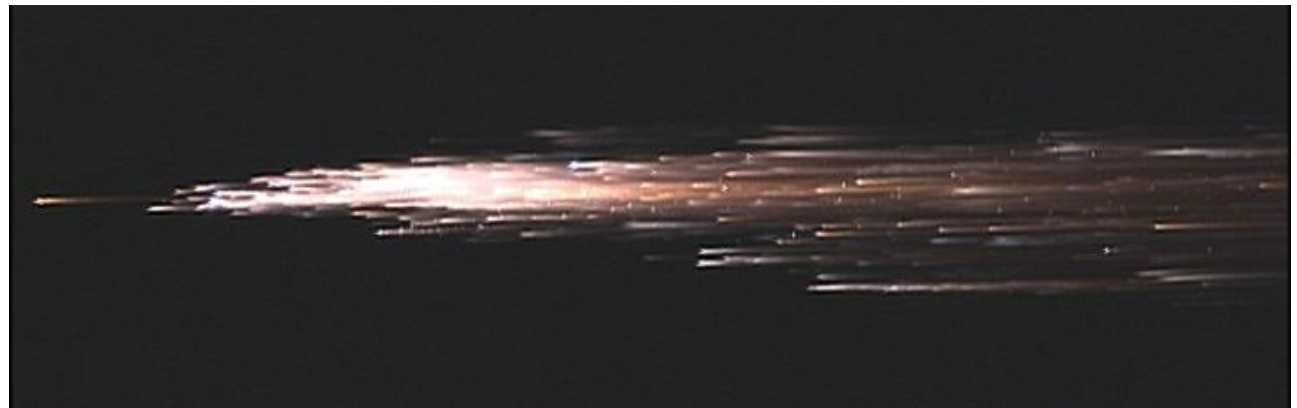
## Reentry of the Jules Verne ATV

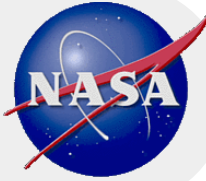
- **NASA and ESA conducted a joint observation campaign of the reentry of the Jules Verne ATV on 29 September 2008.**
  - Two aircraft collected a wide variety of data from vantage points over the Pacific Ocean near the reentry path of the Jules Verne.



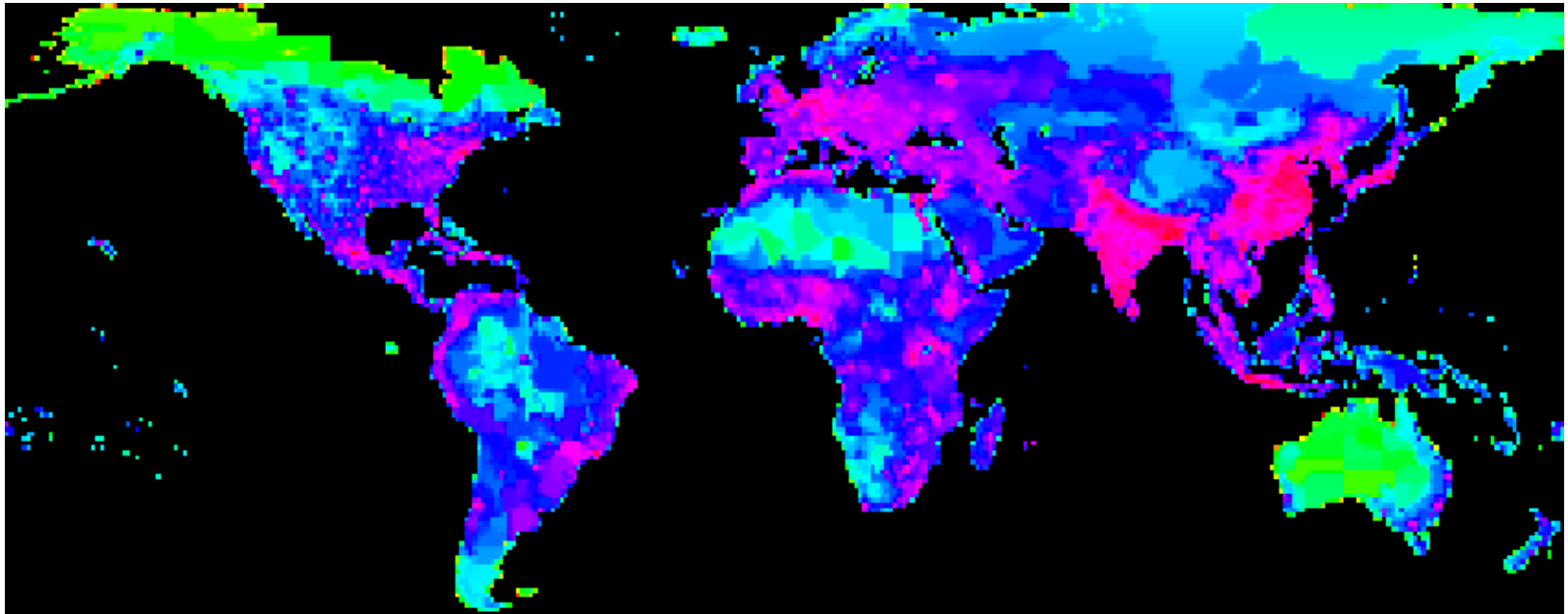
**Jules Verne undocking on  
5 September 2008**

**Reentry over  
Pacific Ocean**





# Population Distribution on the Earth

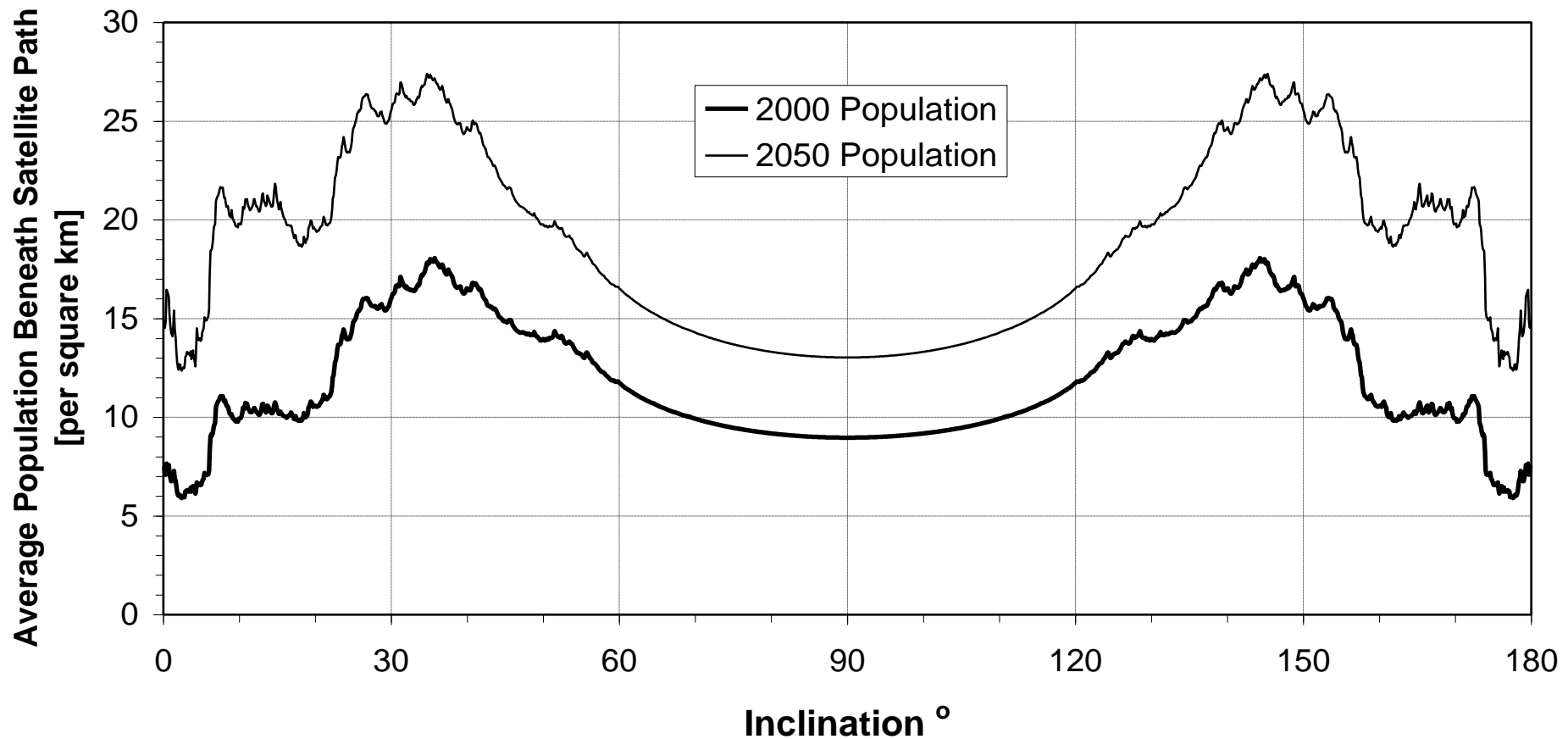


- Gridded Population of the World, version 3 (GPWv3)
- Socioeconomic Data and Applications Center (SEDAC) at Columbia University
- $2.5 \times 2.5$  arc minute cells =  $4.6 \text{ km} \times 4.6 \text{ km}$  cells at the Equator
- Reference years 1990-2015 in 5-year intervals



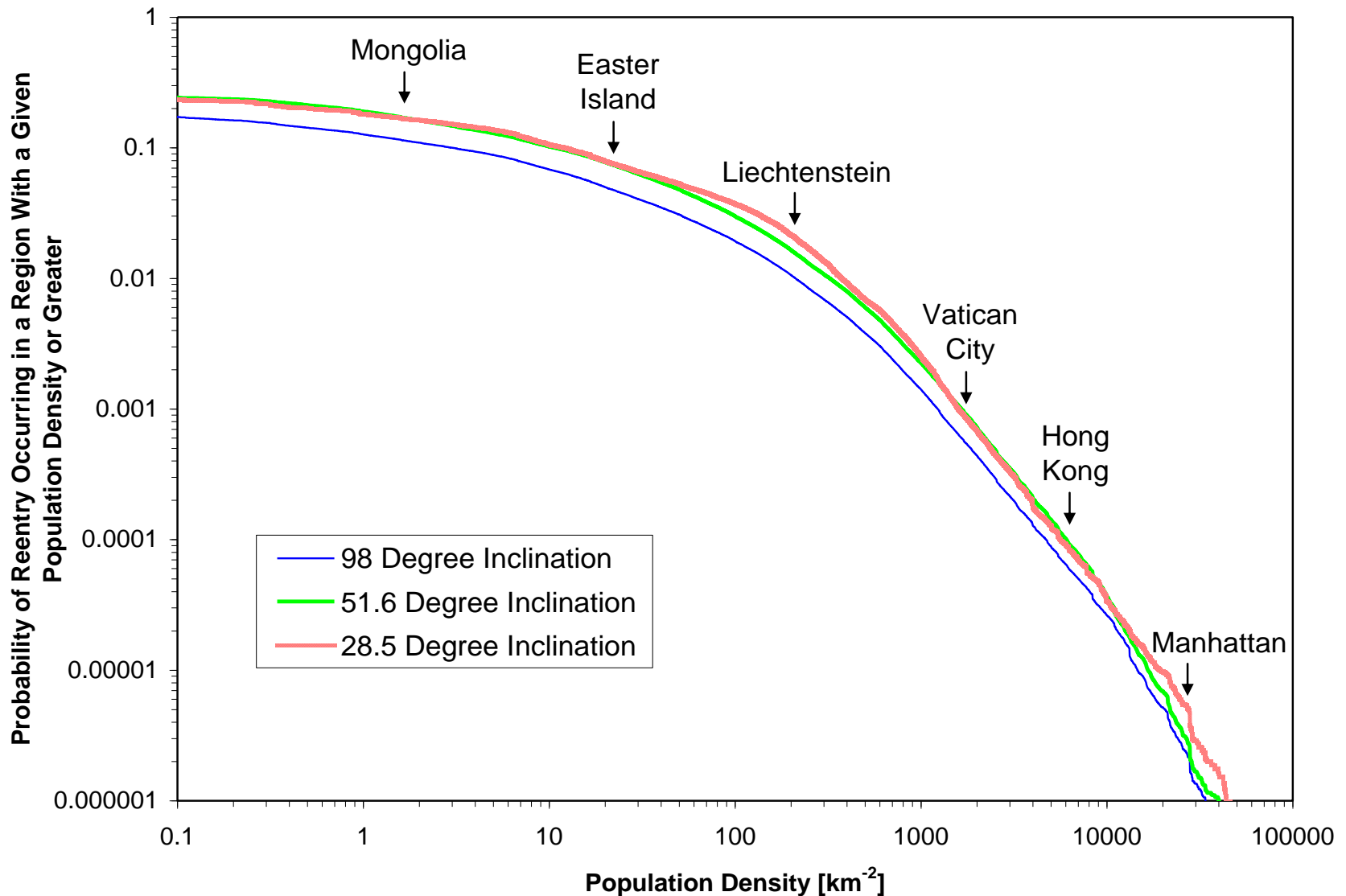
# Average Density of People Below Satellite Path

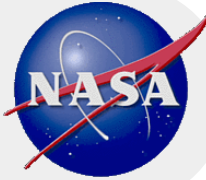
## Inclination-Dependent Latitude-Averaged Population Density





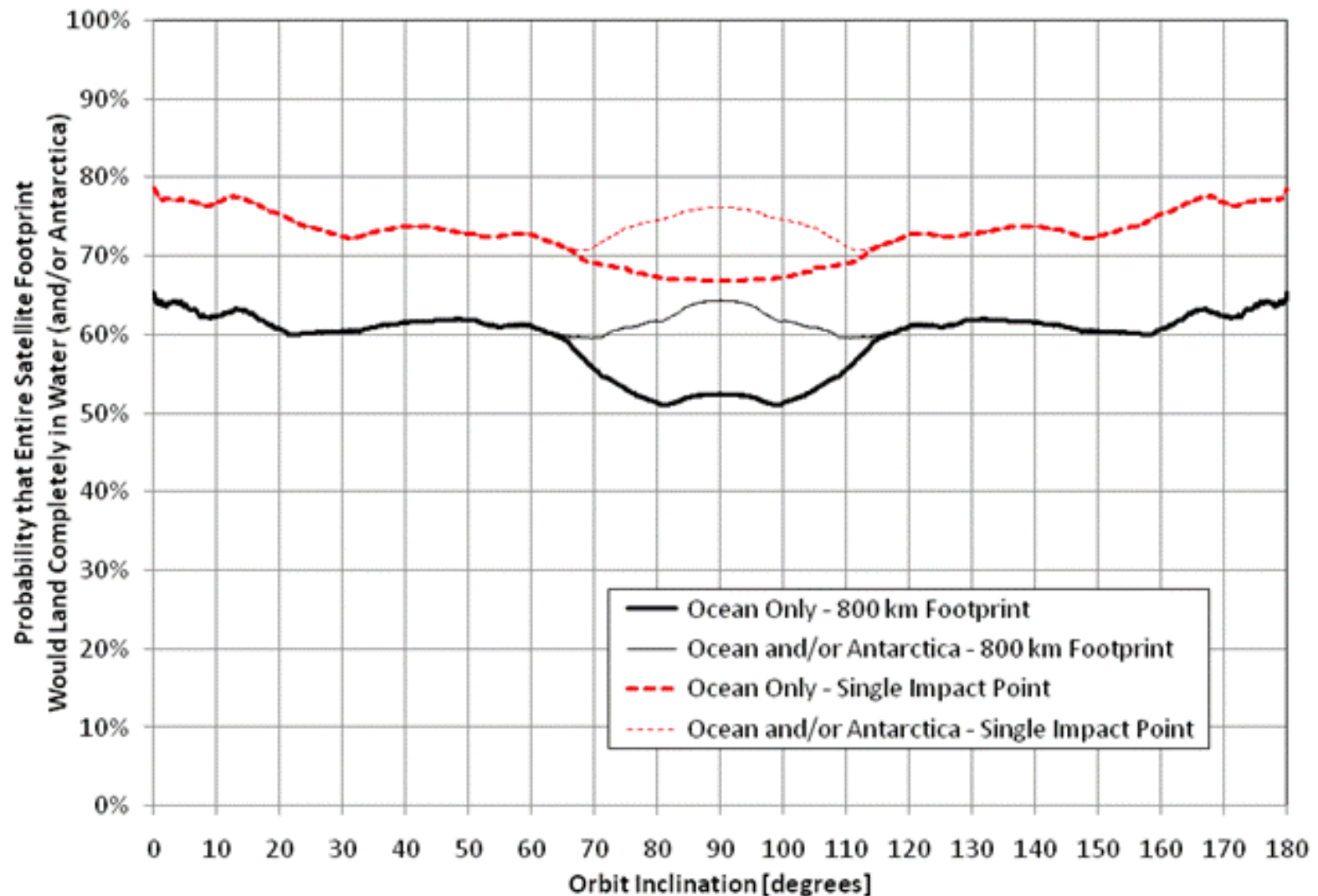
# Probability of Falling in Populated Areas





# Probability of Ocean Reentry

Probabilities of Satellite Reentry Avoiding Land







## Brewster Rokit on Reentry Risks





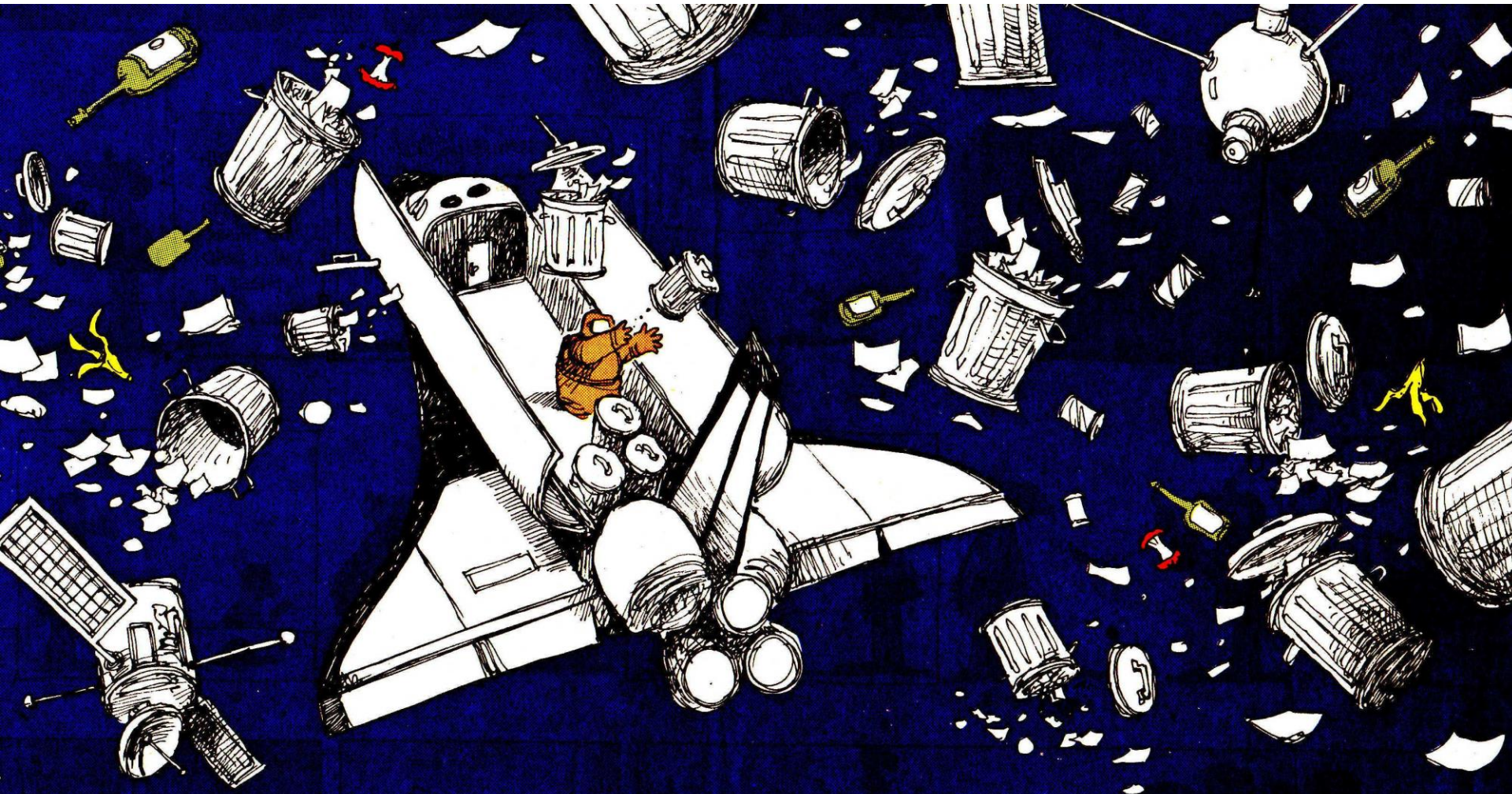
## Conclusions

- **Monitoring the Earth space environment is insufficient to understand the risks posed by debris**
  - Never observe all the debris characteristics you need
  - Cannot see the entire environment
  - Must be able to make predictions about the future
- **Modeling permeates all aspects of orbital debris studies**
- **Models provide users with tools to design their spacecraft to survive the debris environment**
- **Models provide policy makers with tools to be able to make informed decisions about guidelines and regulations concerning space activities**
- **Models are only as good as the assumptions made and the quality of the data behind them**





# Questions?



HOW TO TELL WHEN MAN HAS OFFICIALLY CONQUERED SPACE



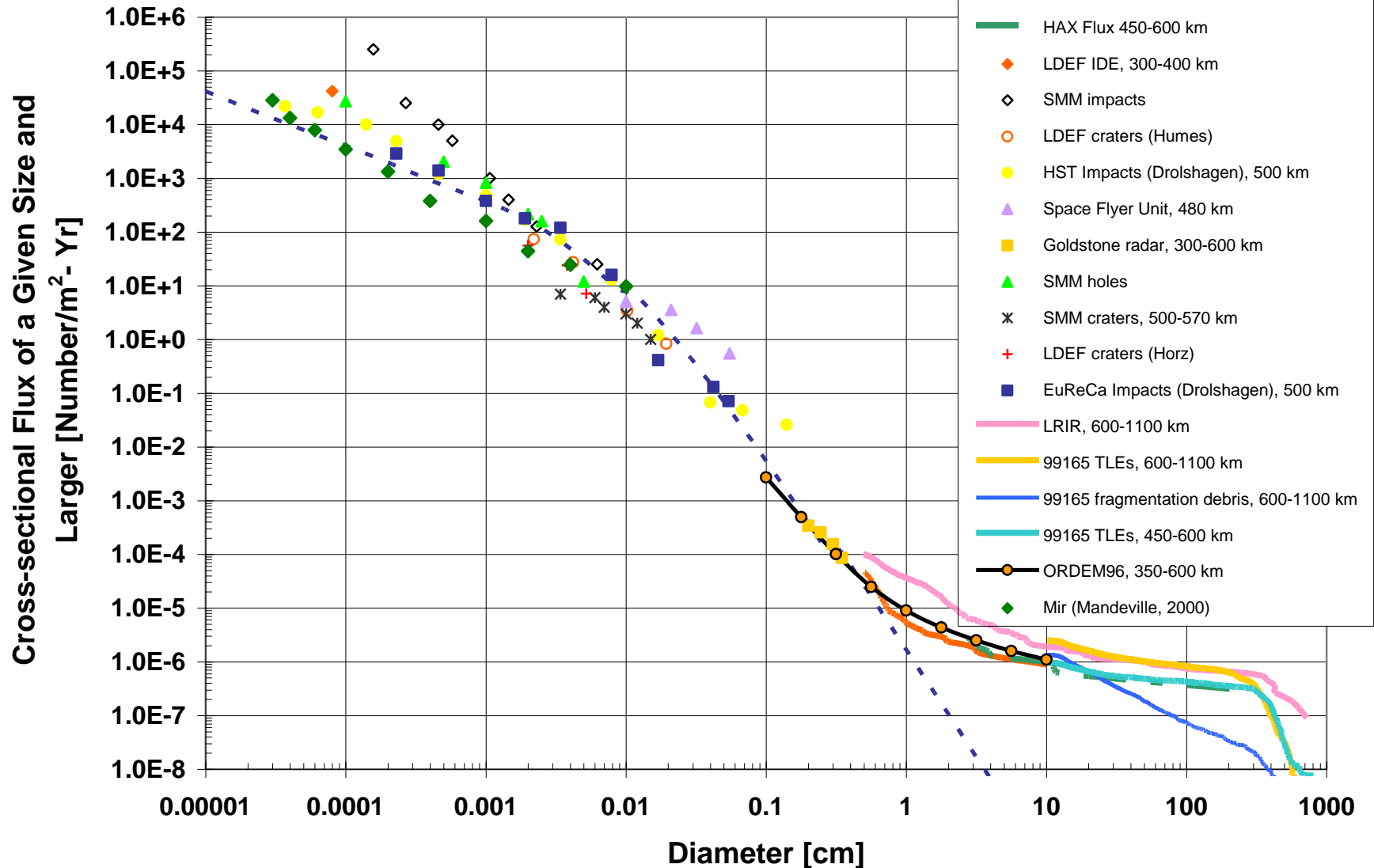
# Backup





# Data Compilation

## Orbital Debris Environment





## Breakup Model

- **NASA's Breakup Model can be used to simulate the evolution of individual breakups**
- **On August 6, 2012, the Russians attempted to launch two communications satellites using a Proton rocket**
- **The BRIZ-M upper stage failed to burn properly, and was left stranded in an elliptical orbit with about 5 metric tons of its propellant still aboard**
- **On October 16, the rocket body exploded, creating at least 700 trackable pieces of debris (and probably many more too small to be tracked) in orbits that cross ISS altitude**
- **Observed by astronomers at the Siding Springs Observatory**







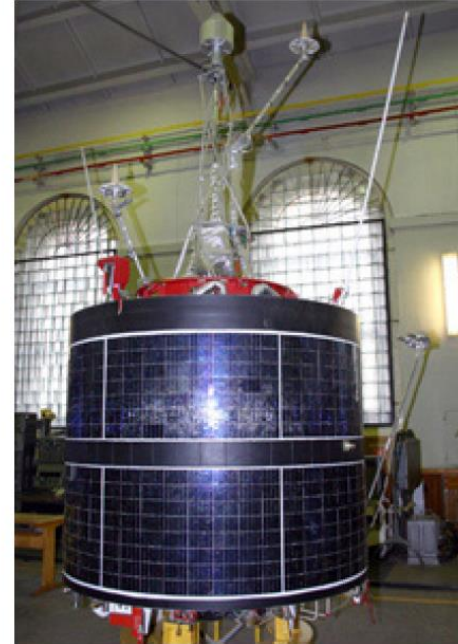
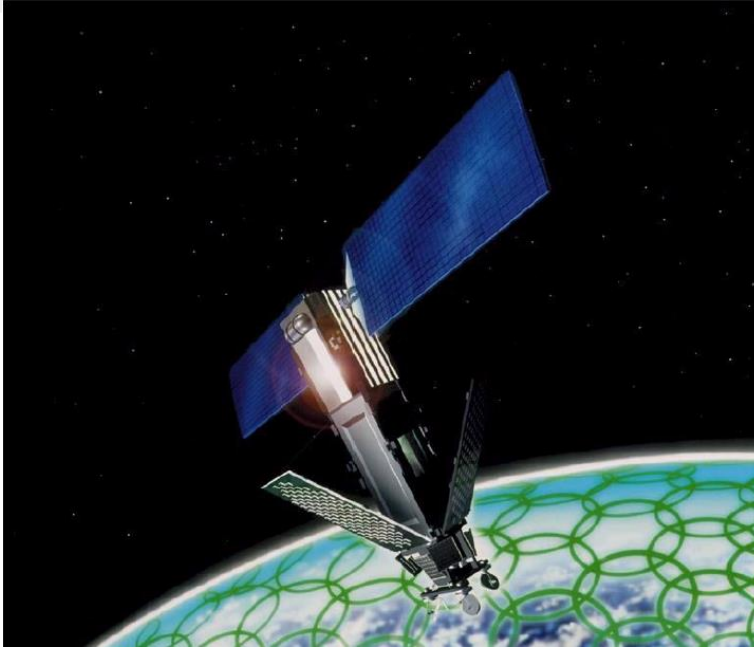
# BRIZ-M Explosion



2012/10/16 14:05:00 UT



## 2009 Collision



February 10, 16:56 GMT two satellites collided near 789 km altitude

Iridium 33 (24946, 97051C)

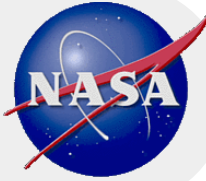
779 x 808 km, 86.4° orbit, 556 kg

Operational US Commercial Communication Satellite

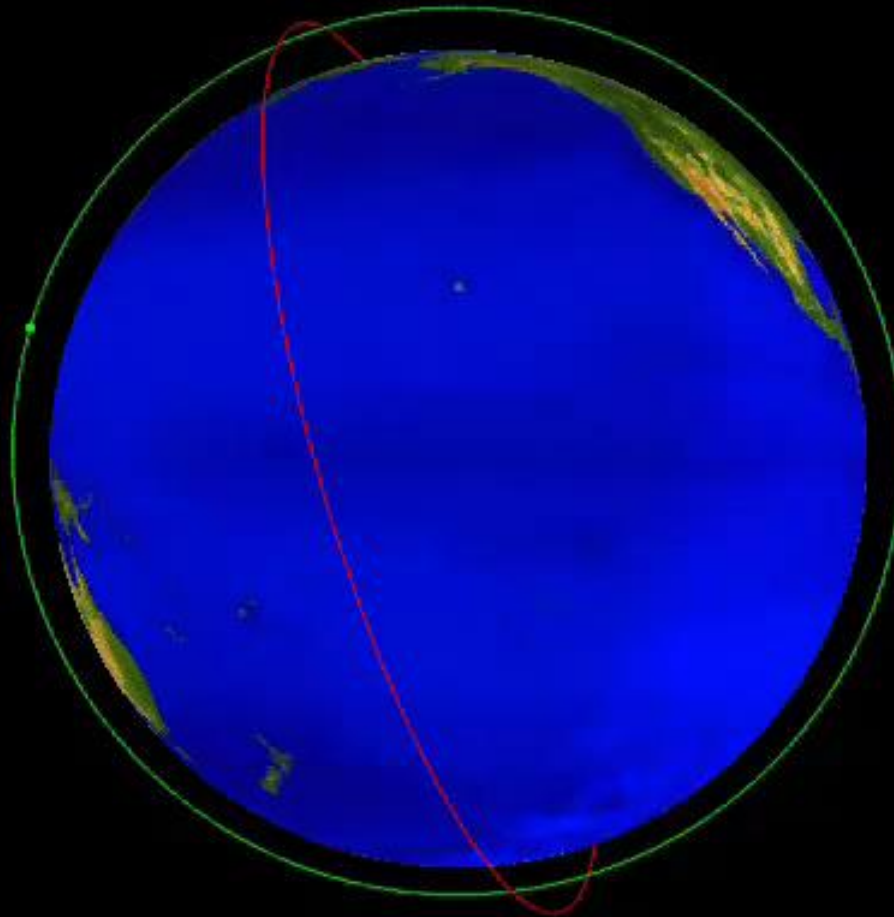
Kosmos 2251 (22675, 93036A)

786 x 826 km, 74.0° orbit, 900 kg

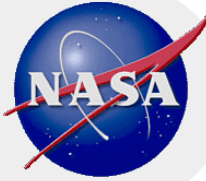
Non-operational Russian Communication Satellite



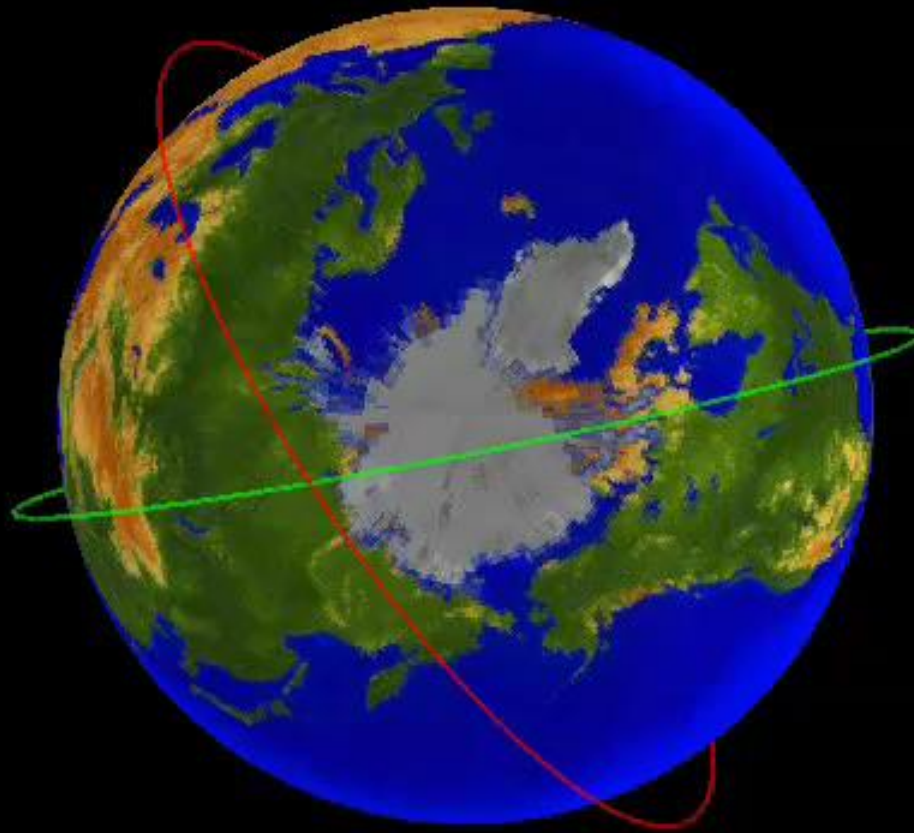
## 2009 Collision



2009/02/10 15:00:00 UT



## 2009 Collision



2009/01/15 18:38:16 UT